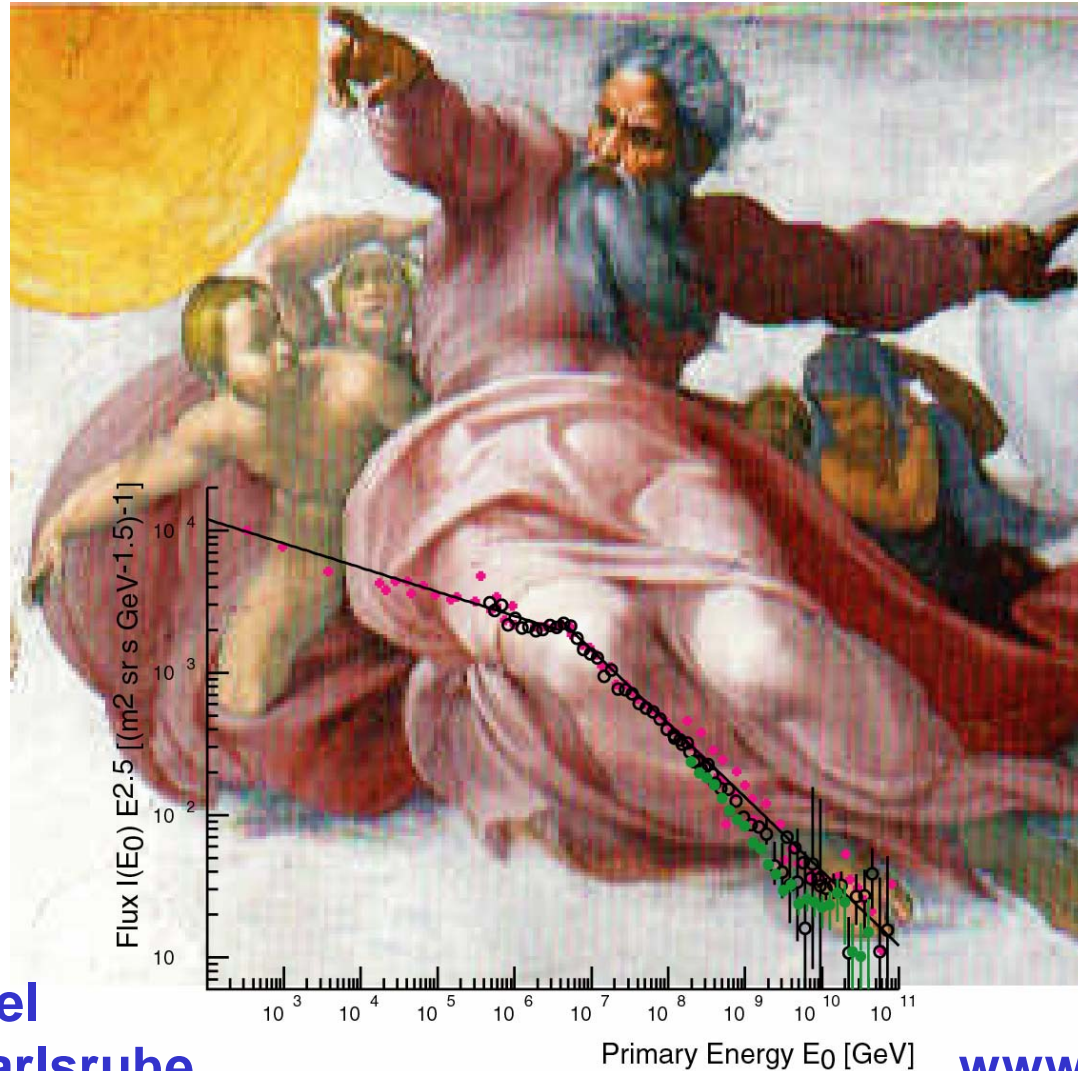


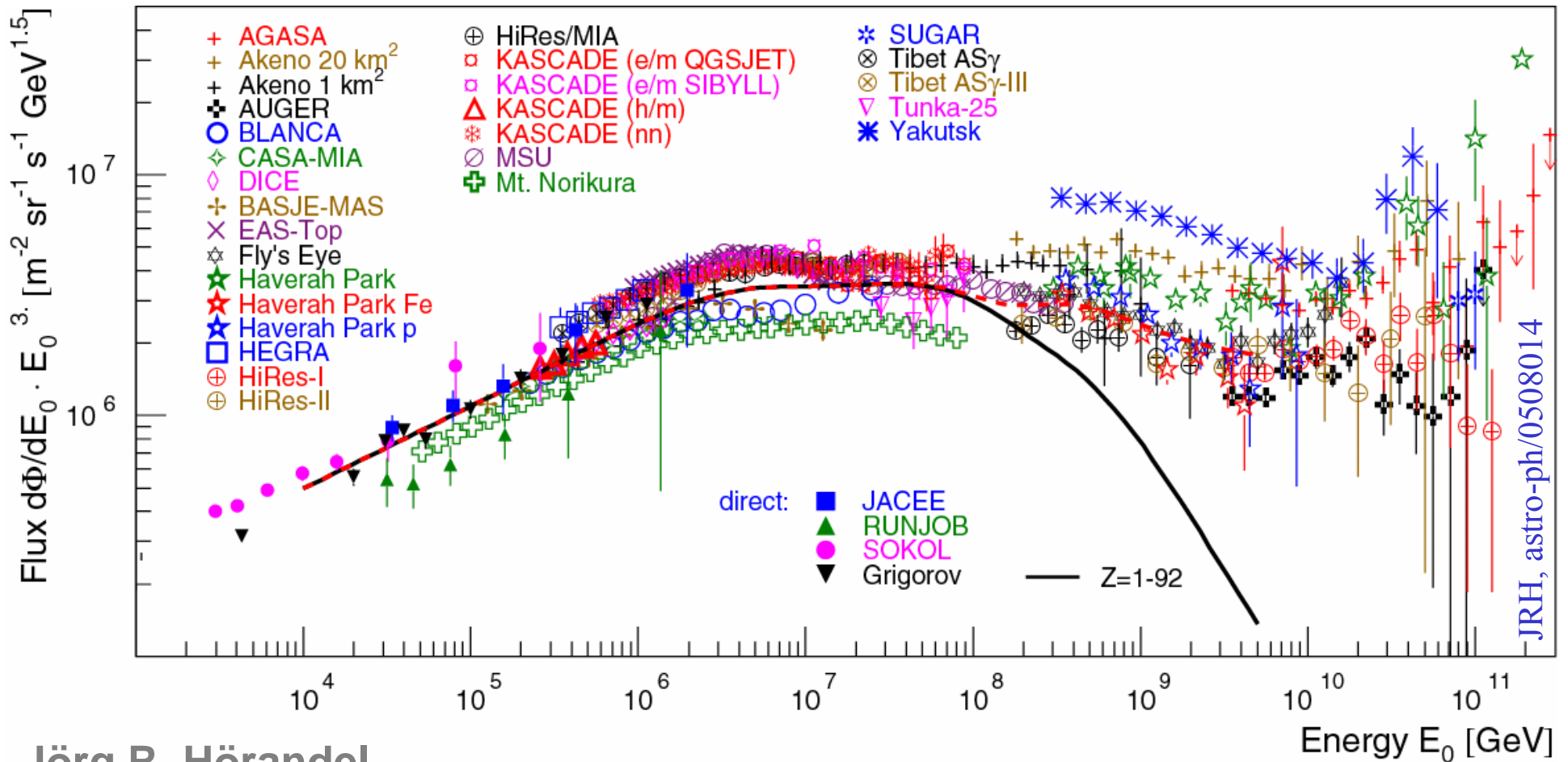
From the Knee to the toes: The challenge of cosmic-ray composition



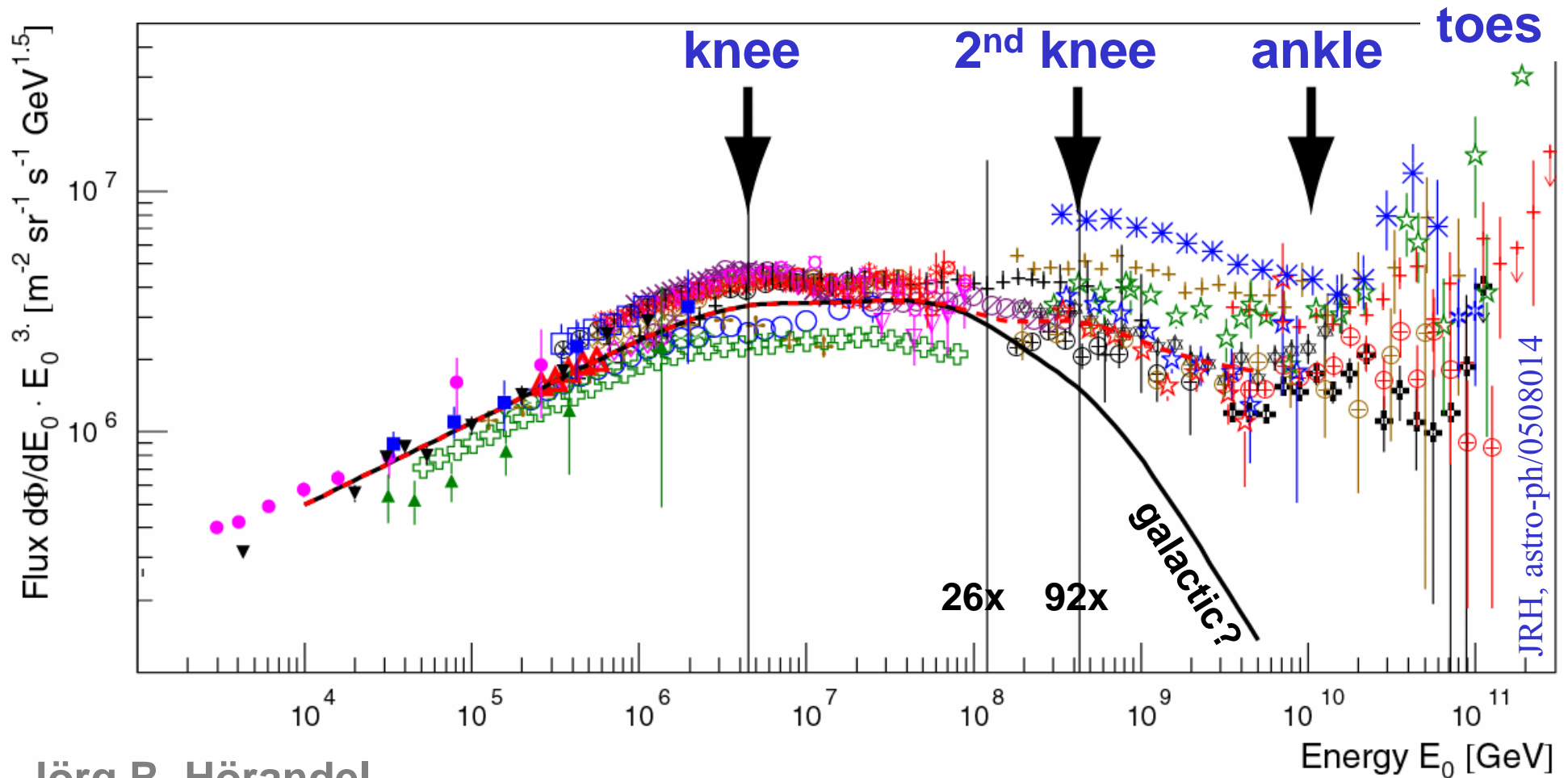
Jörg R. Hörandel
University of Karlsruhe

www-ik.fzk.de/~joerg

From the Knee to the toes: The challenge of cosmic-ray composition



From the Knee to the toes: The challenge of cosmic-ray composition



Jörg R. Hörandel
University of Karlsruhe

www-ik.fzk.de/~joerg



International Cosmic-Ray Conference, Chicago, 1939



1. A. H. Compton
 2. A. Biñls
 3. M. S. Vallarta
 4. C. Anderson
 5. V. F. Hess
 6. W. Bothe
 7. W. Heisenberg
 8. J. Clay

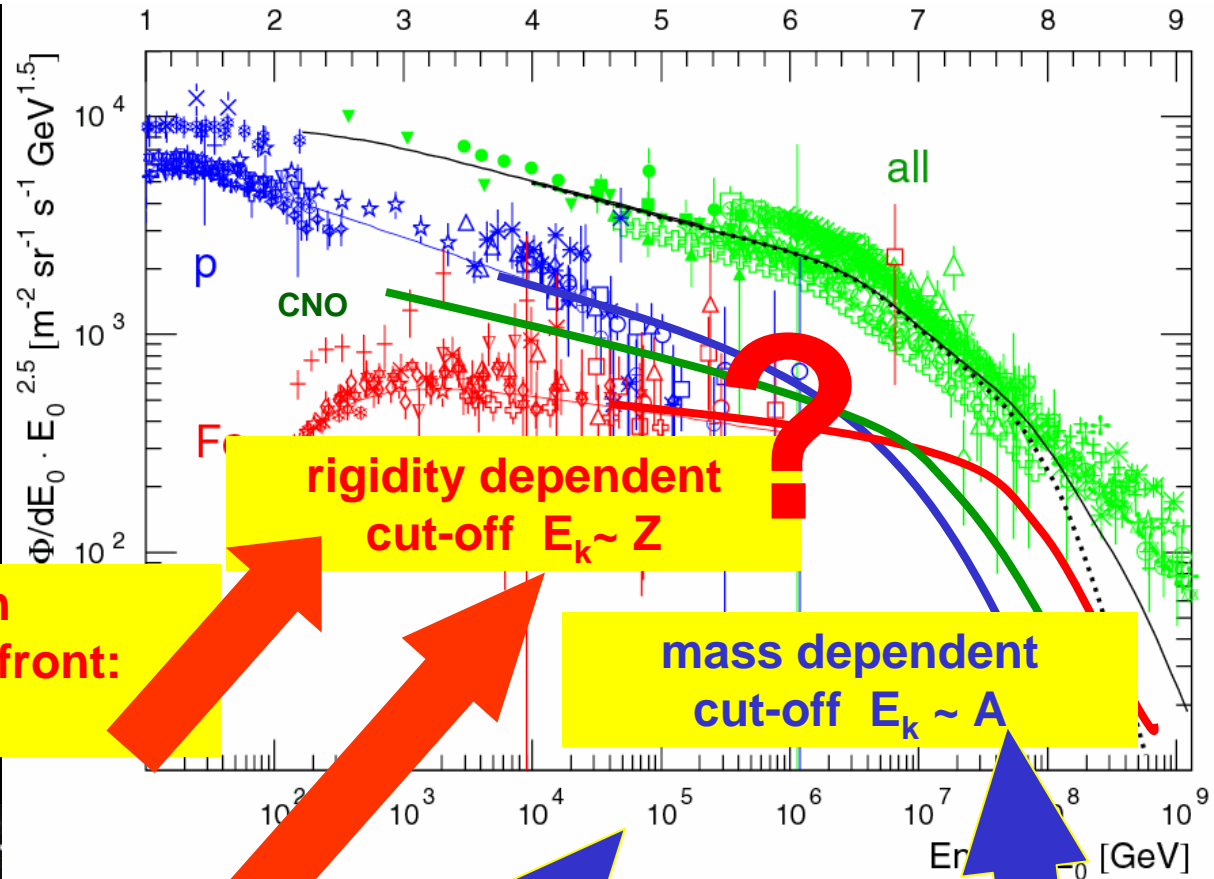
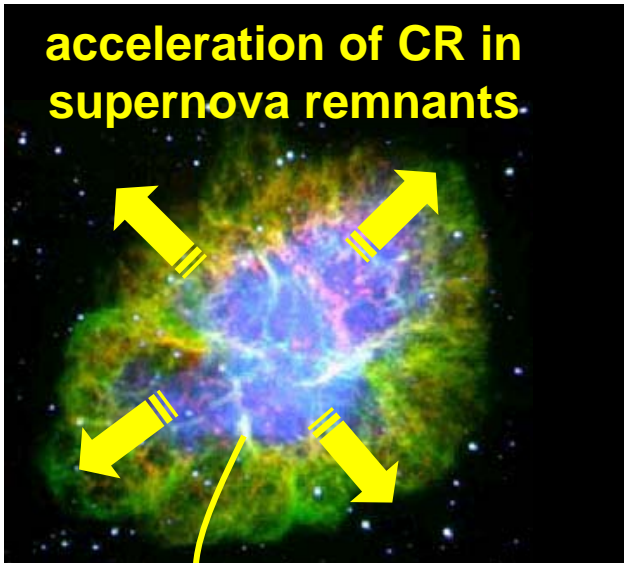
9. W. F. G. Swann
 10. G. Herzog
 11. Beutler
 12. W. D. Harkins
 13. S. Neddermeyer
 14. T. H. Johnson
 15. P. Auger
 16. B. Rossi

17. J. R. Oppenheimer
 18. S. Goudsmit
 19. E. Teller
 20. R. Brode
 21. H. Bethe
 22. C.G. Montgomery
 23. W. Bostick
 24. L. Nordheim

25. S.E. Forbush
 26. V. Wilson
 27. J. Wheeler
 28. M. Pomerantz
 29. Dershem
 30. R. Serber
 31. E. Wollan
 32. M. M. Shapiro

33. C. Eckart
 34. D. Hughes
 35. W. Jesse
 36. J.B. Hoag
 37. N. Hilberry
 38. P. Gilli
 39. A. Snett

acceleration of CR in supernova remnants



Fermi acceleration
finite lifetime of shock front:
 $E_{\text{max}} \sim Z \cdot 10^{15} \text{ eV}$

rigidity dependent cut-off $E_k \sim Z$

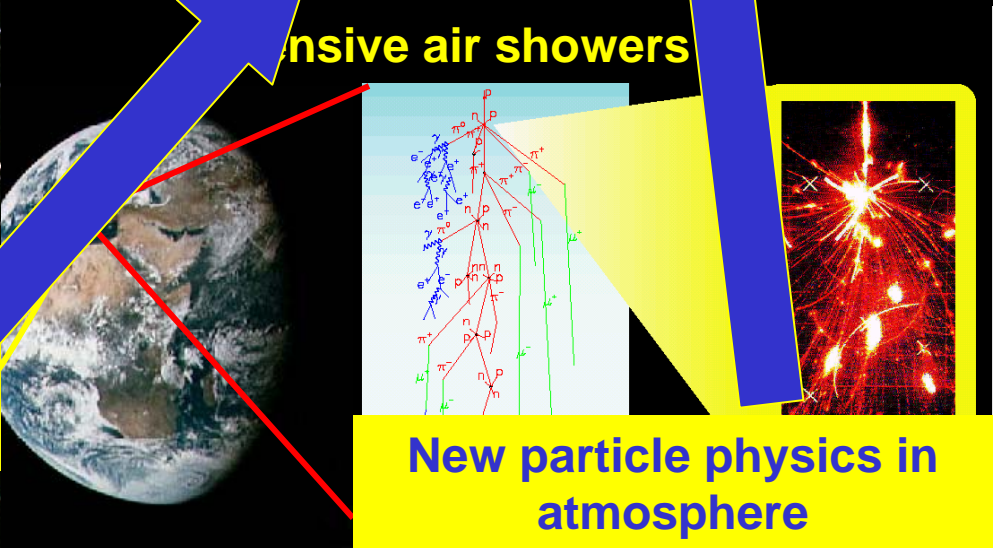
mass dependent cut-off $E_k \sim A$

propagation through galaxy

Leakage from Galaxy:
escape probability $\sim f(Z)$

$B = 3 \mu\text{G}$

Interactions with background particles (photons, neutrinos)



extensive air showers

New particle physics in atmosphere

TRACER Experiment - Mc Murdo, Antarctica

LDB flight: December 12th – 26th, 2003

~ 40 km (3-5 g/cm²)

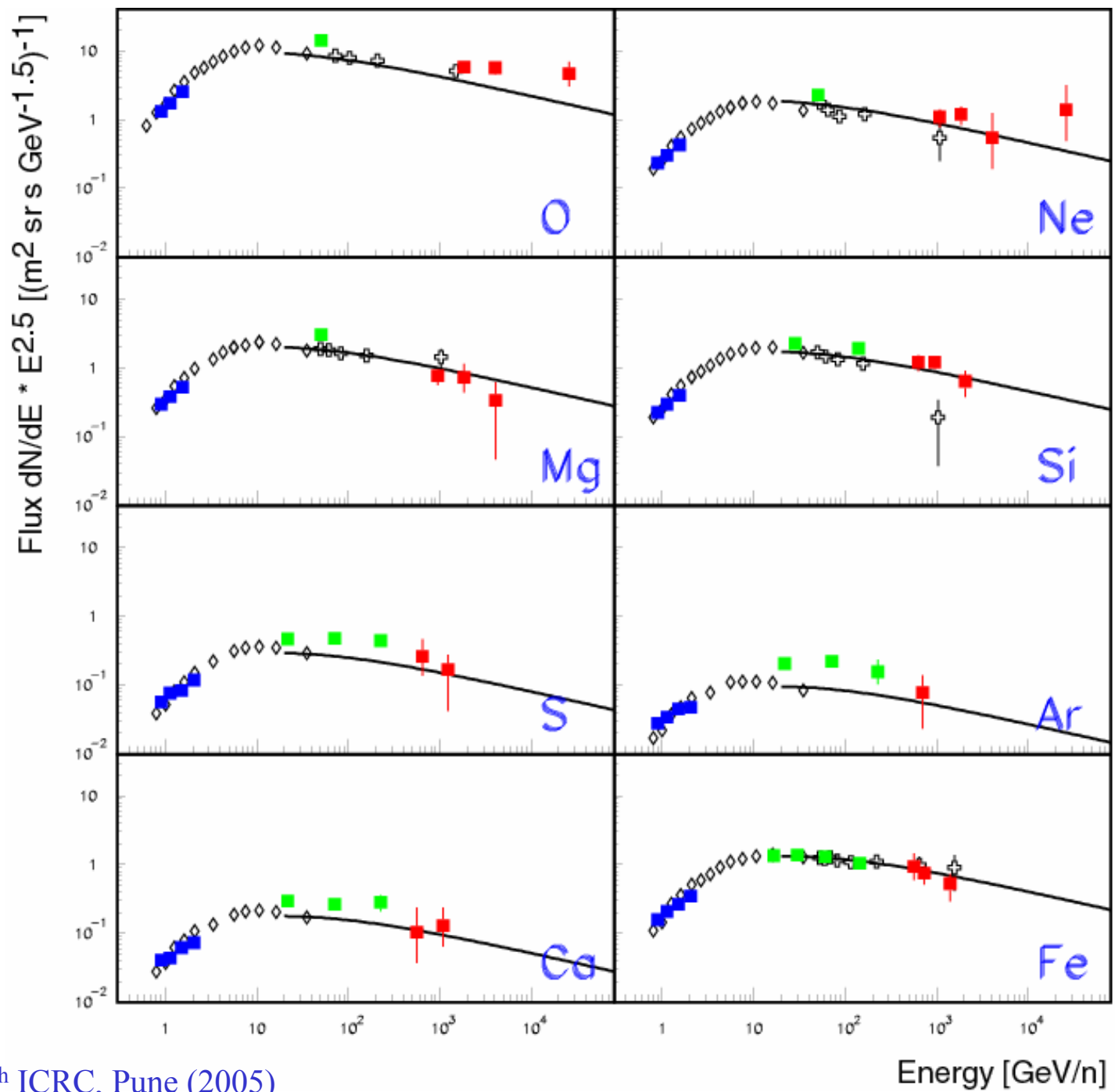


Z = 8 - 26

E = 1 GeV/n - ~10 TeV/n



TRACER energy spectra for individual elements



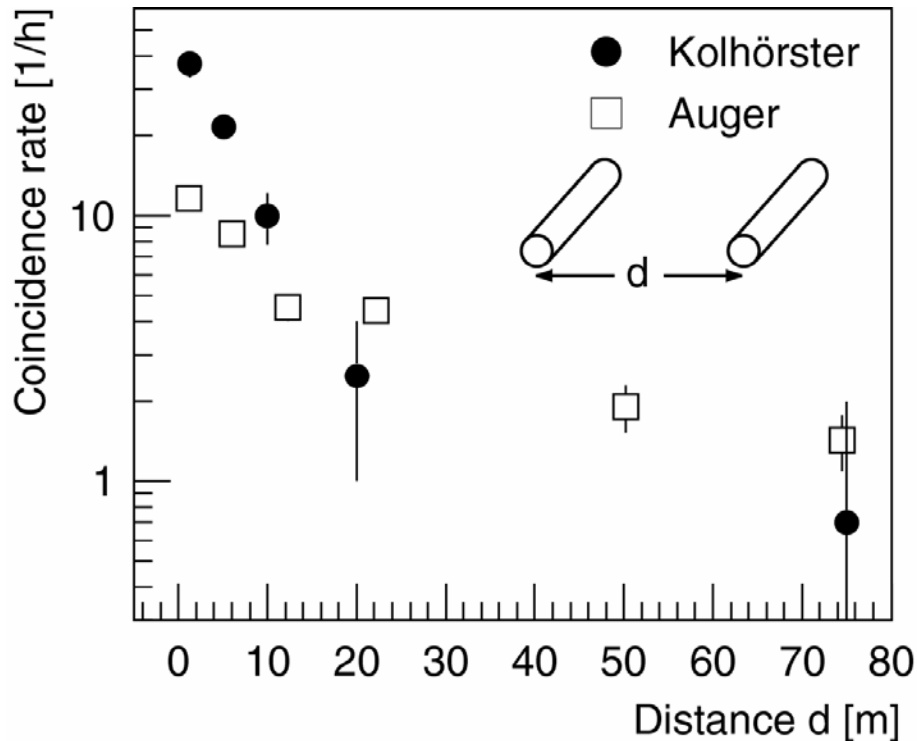
Extensive air showers



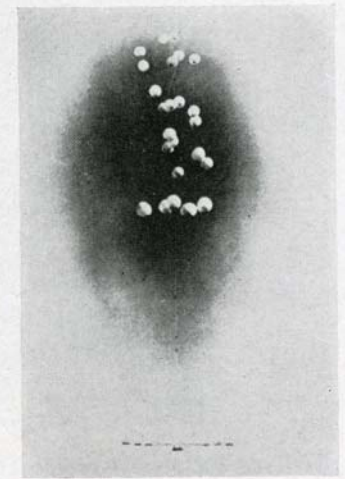
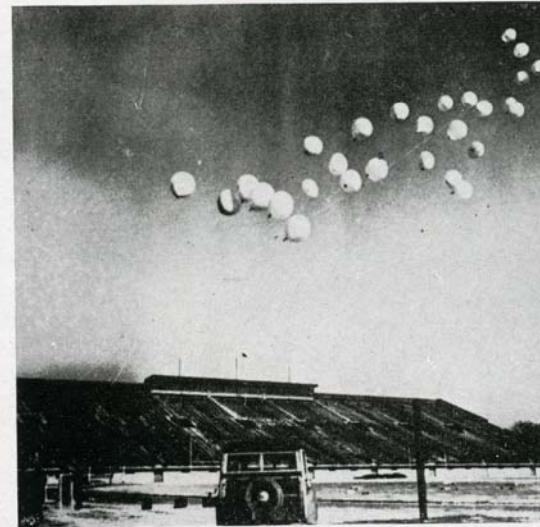
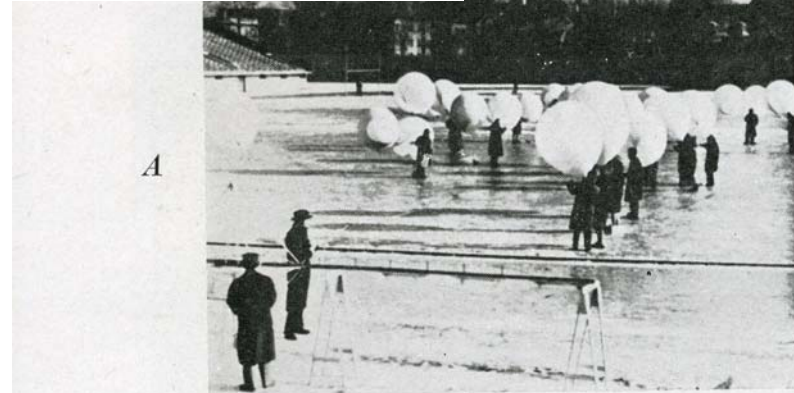
**P. Auger
Jungfrauoch**

MEASURING COSMIC RAYS IN THE SWISS ALPS

The author (*left*) and his collaborator, P. Ehrenfest, set up their apparatus in the Jungfrauoch.



W. Kolhörster et al., *Naturwiss.* 26 (1938) 576
 P. Auger et al., *Comptes renduz* 206 (1938) 1721



P. Auger
 BALLOON FLIGHT OF JANUARY, 1943, CONDUCTED BY THE AUTHOR, SCHEIN,
 AND ROGOZINSKI FOR THE MEASUREMENT OF EXTENSIVE (OR
 AUGER-) SHOWERS IN THE STRATOSPHERE

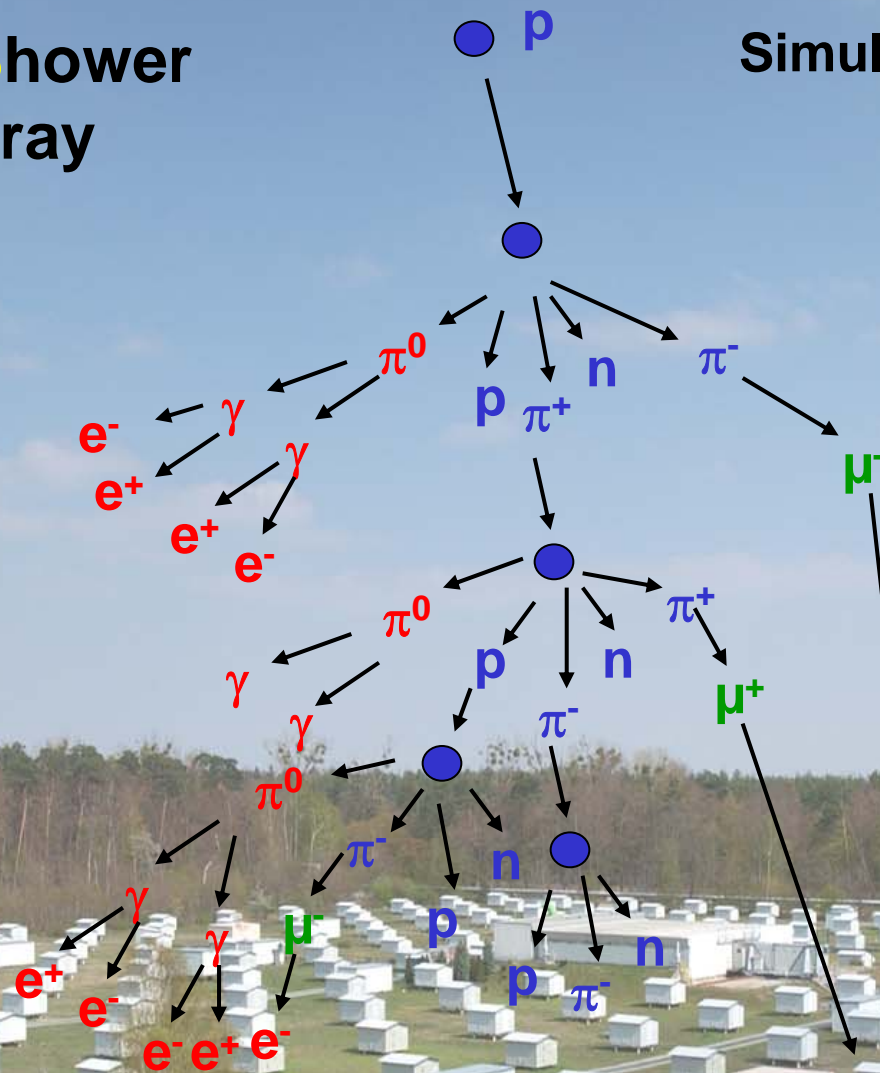
A. The balloons are assembled on Stagg Field at the University of Chicago, Chicago, Illinois. In the foreground can be seen the long frame which was required for the wide separation of the cosmic-ray counters.

B. The large cluster of balloons as it is about to be released.

C. The balloon train sails into the sky after its release. Suspended below the balloons is the frame supporting the counters and recording apparatus.

KARlsruhe Shower Core and Array DEtector

Simultaneous measurement of
electromagnetic,
muonic,
hadronic
shower components

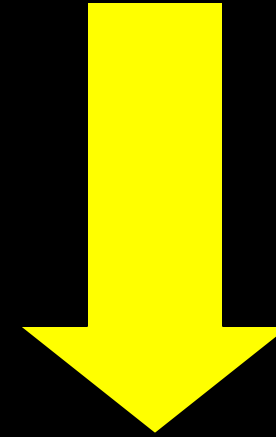


acceleration of CR in
supernova remnants

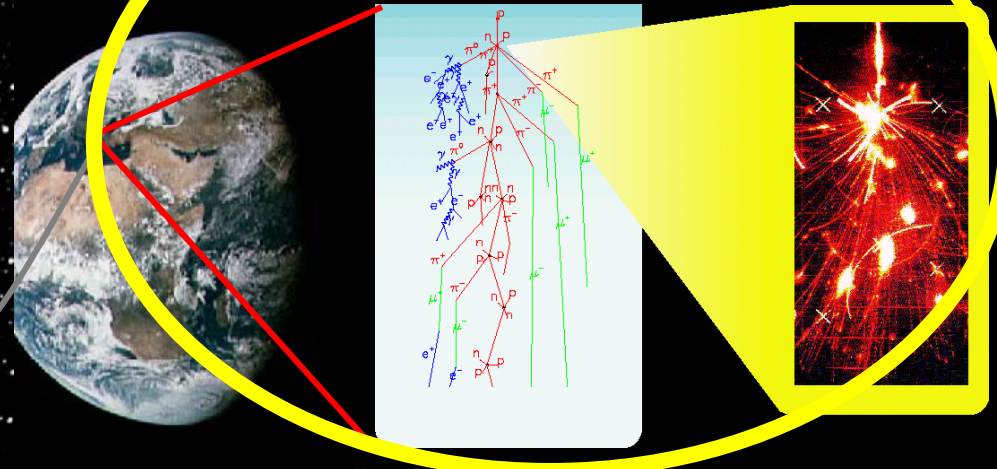
Interactions

propagation through
galaxy

$B = 3 \mu\text{G}$



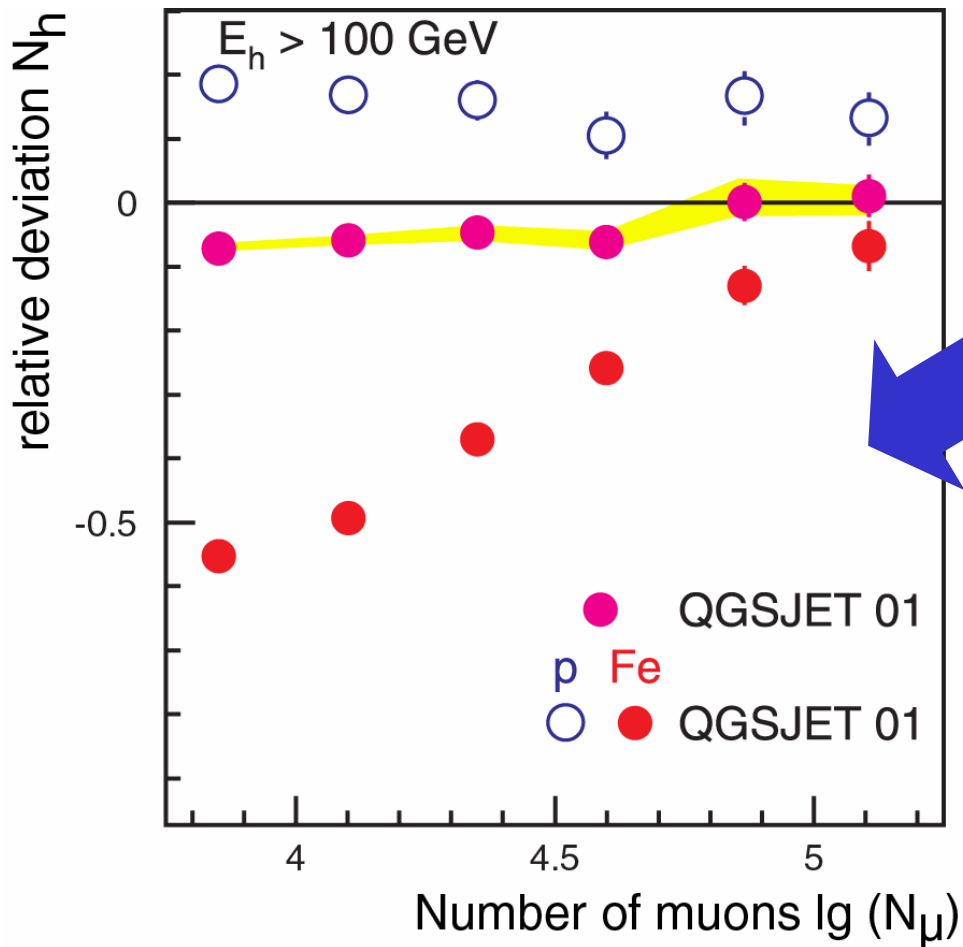
extensive air showers



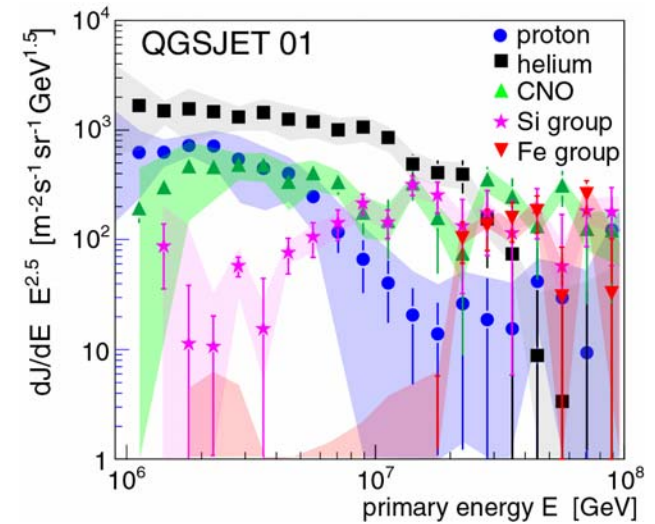
KASCADE: Test of interaction models

QGSJET 01

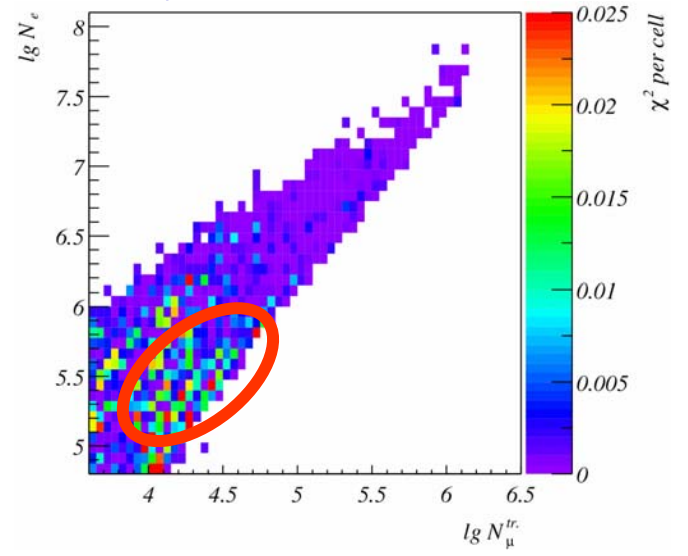
Number of hadrons vs. number of muons



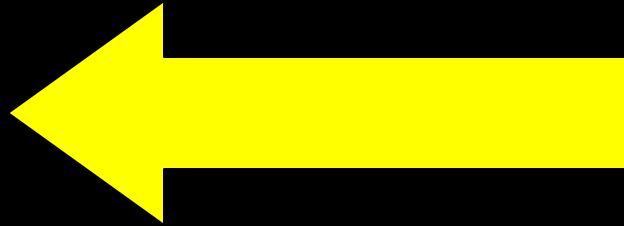
N_e - N_μ analysis



χ^2 distribution

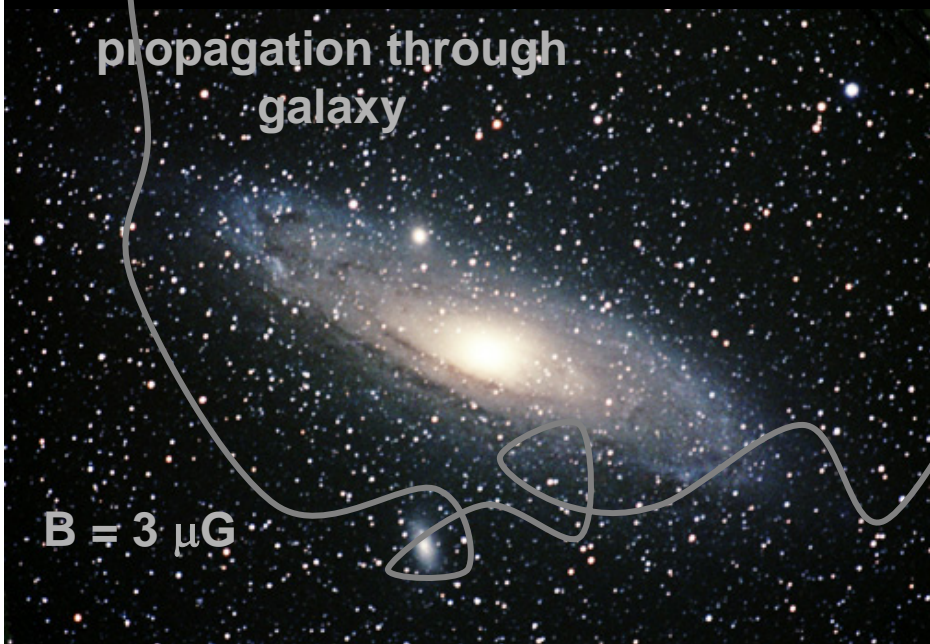


acceleration of CR in
supernova remnants



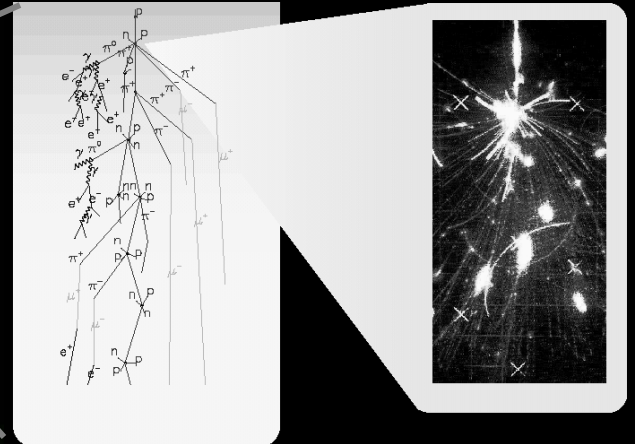
Sources

propagation through
galaxy

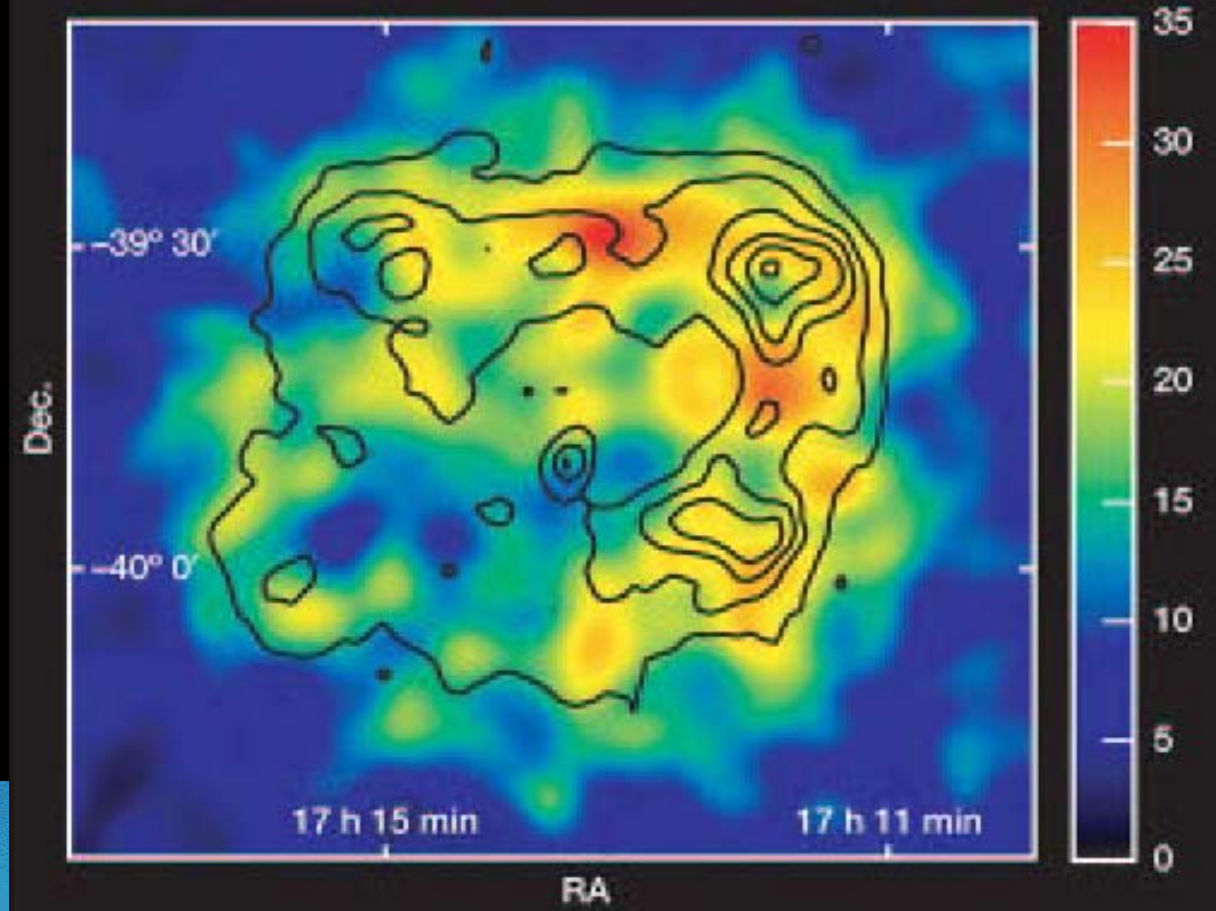
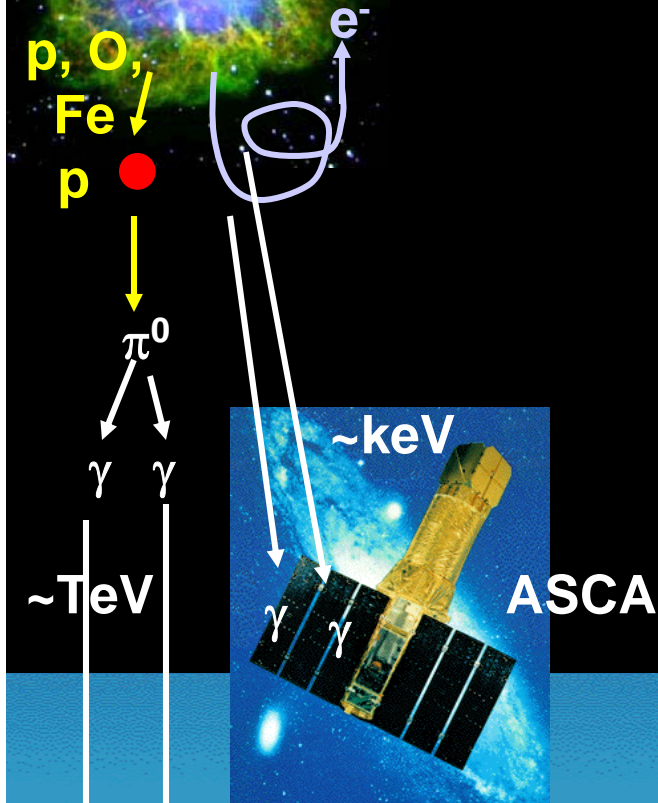


$B = 3 \mu\text{G}$

extensive air showers



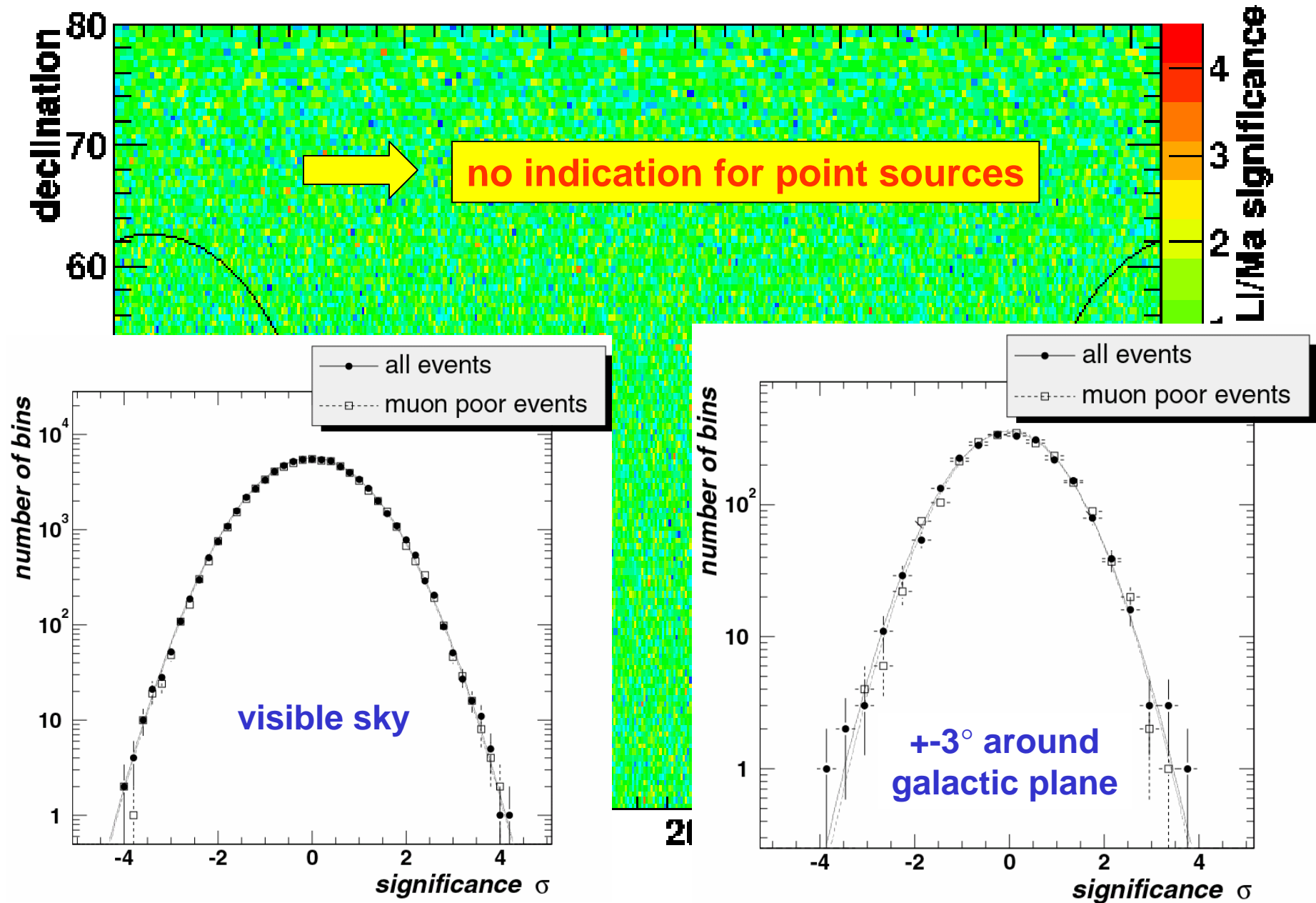
Acceleration of particles in supernova remnant



SN R RX J1713.7-3946
H.E.S.S.: TeV-Gamma rays
ASCA: X-rays (keV)

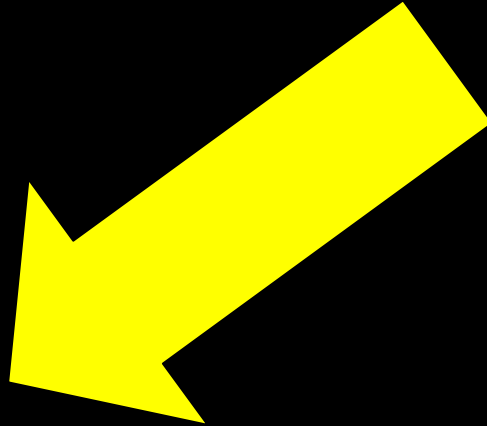


KASCADE: Small scale anisotropy – point source search



acceleration of CR in
supernova remnants

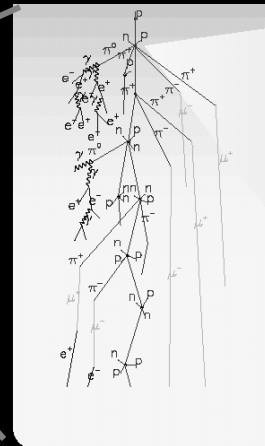
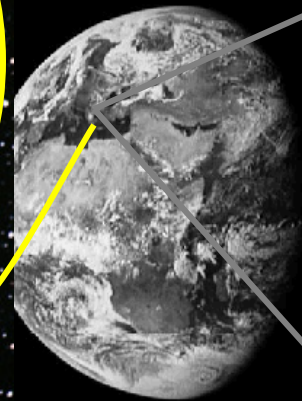
Propagation



propagation through
galaxy

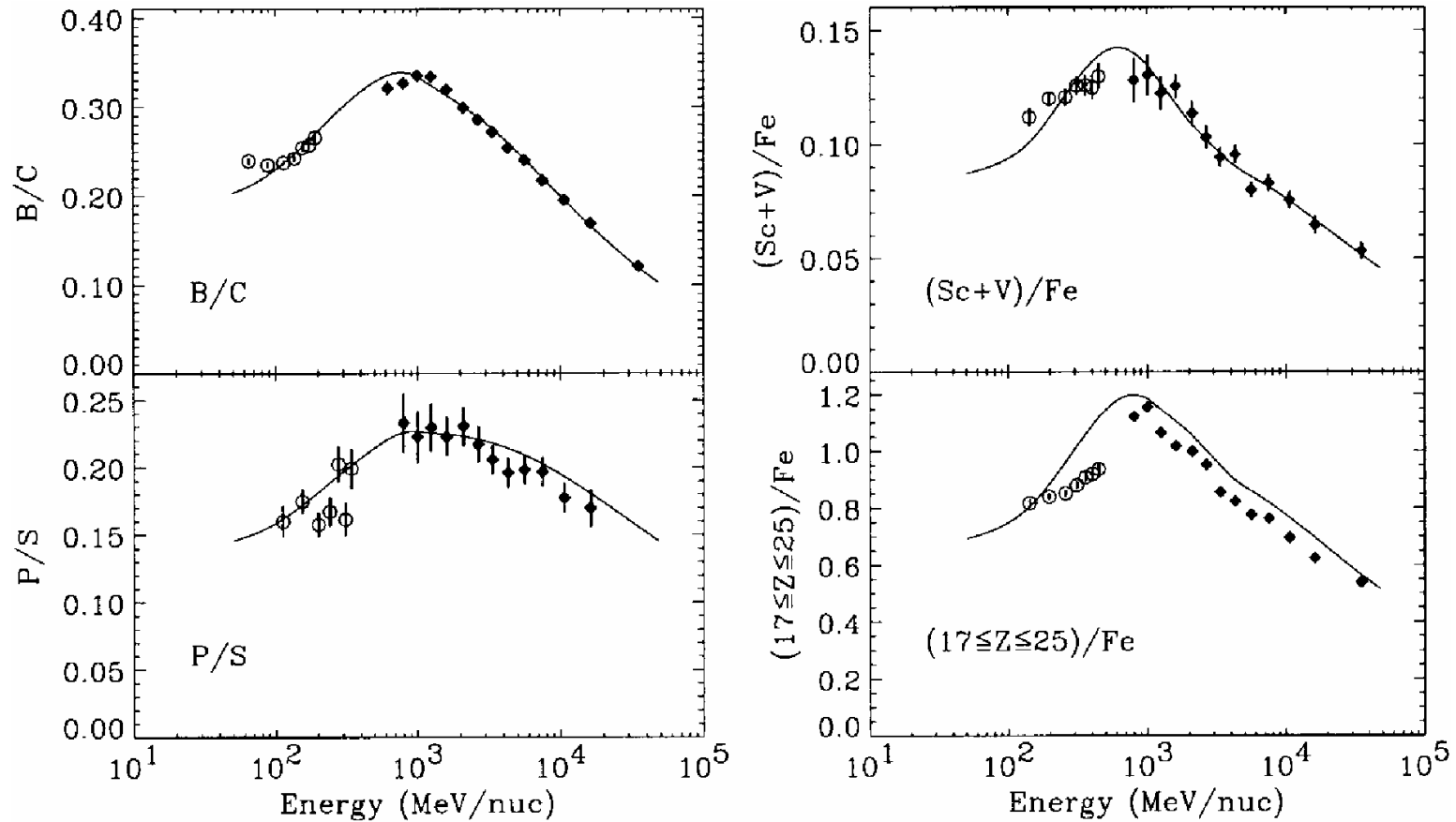
$B = 3 \mu\text{G}$

extensive air showers



Ratio of secondary to primary nuclei

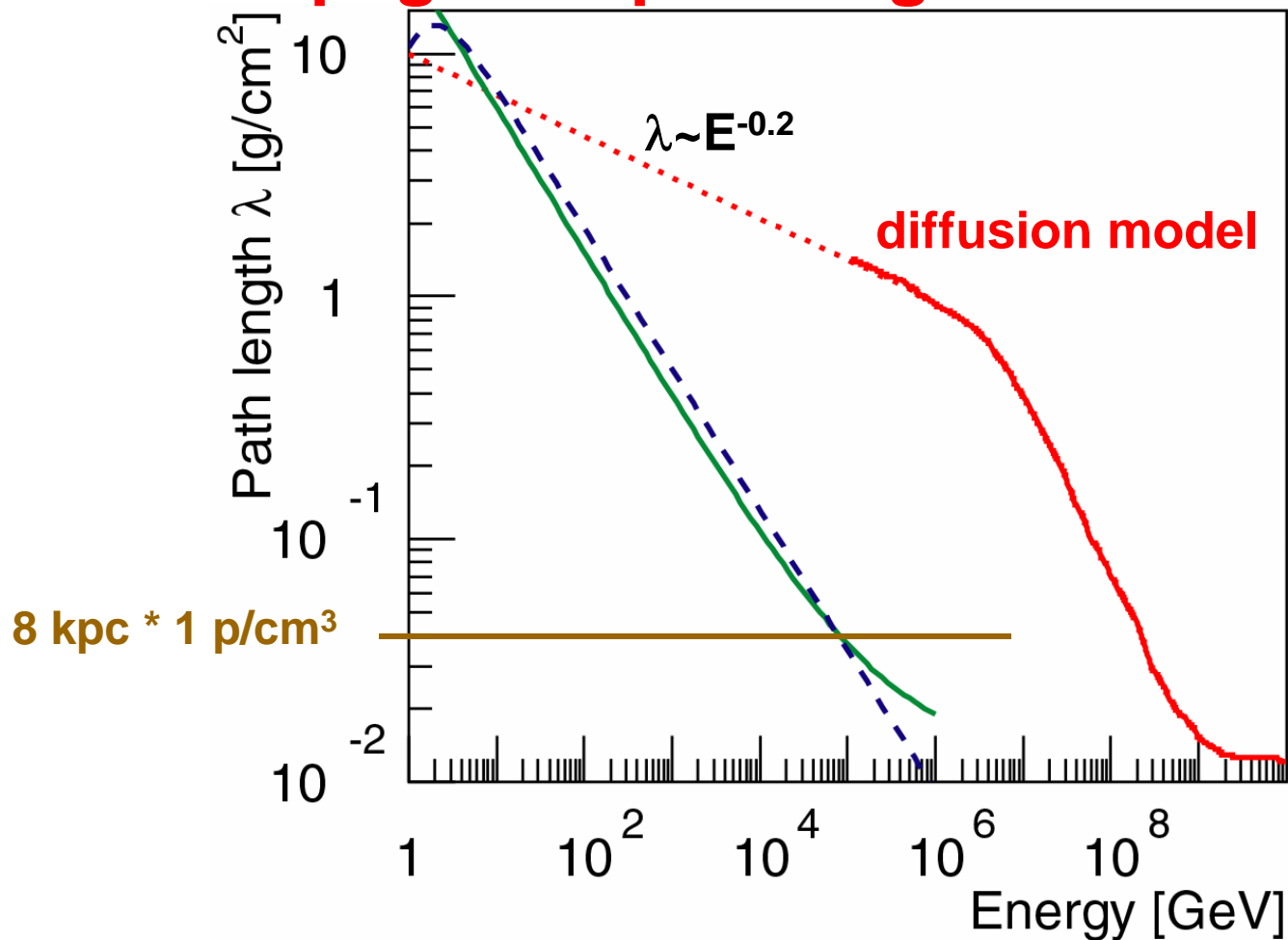
ACE/CRIS



Leaky box model

$$\lambda_{\text{esc}} = \frac{26.7 \beta \text{g/cm}^2}{(\beta R / 1.0 \text{ GV})^{0.58} + (\beta R / 1.4 \text{ GV})^{-1.4}}$$

Propagation pathlength in Galaxy



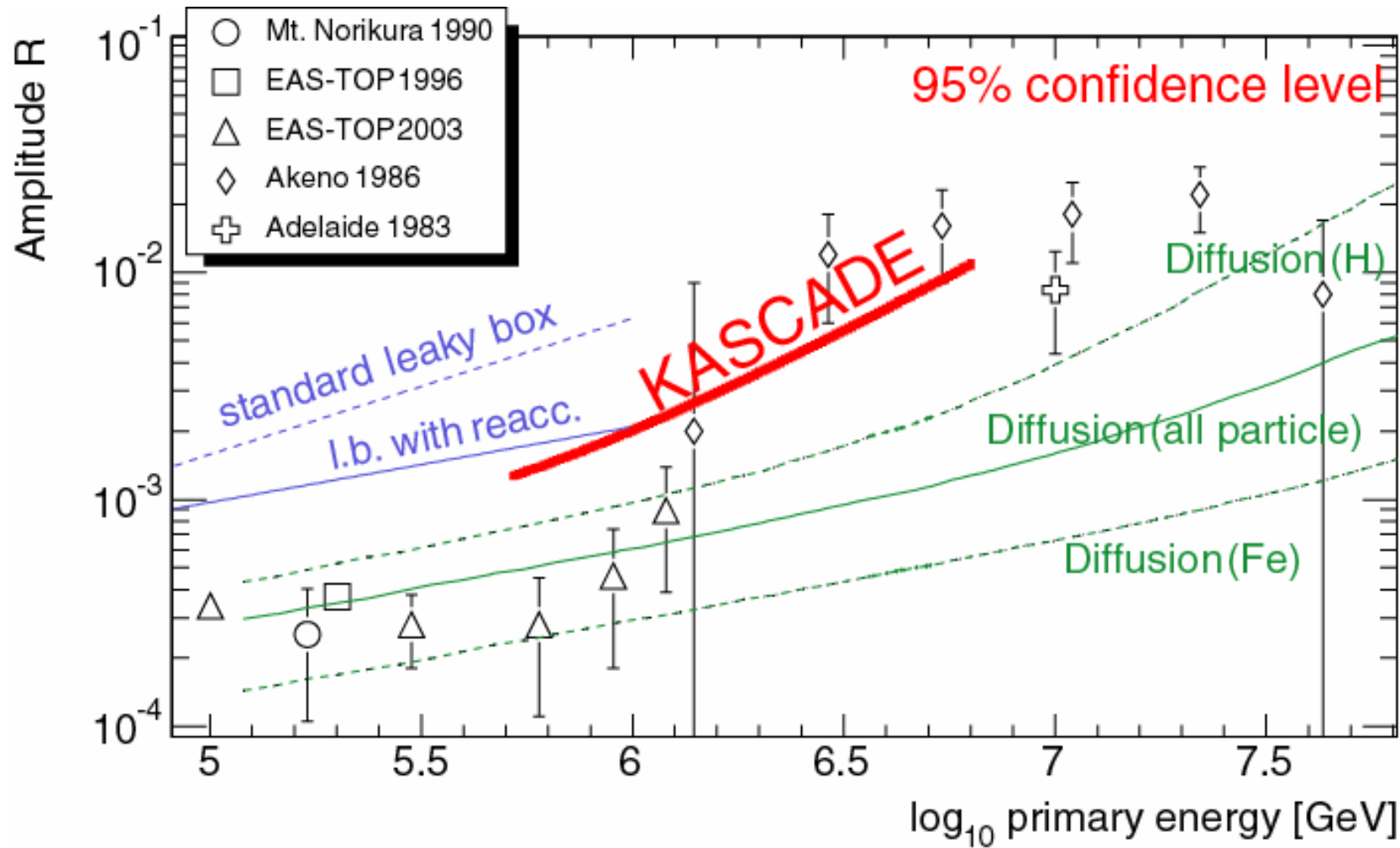
leaky box model

$$\lambda = \frac{26.7 \beta \text{g}/\text{cm}^2}{(\beta R / 1.0 \text{ GV})^{0.58} + (\beta R / 1.4 \text{ GV})^{-1.4}}$$

residual pathlength model

$$\lambda = \left[6.0 * \left(\frac{R}{10 \text{ GV}} \right)^{-0.6} + 0.013 \right] \text{g}/\text{cm}^2$$

Anisotropy amplitude vs energy



Rayleigh vector

$$|\vec{R}| = \left(\frac{2}{n} \sum_{i=1}^n \sin \alpha_i \right)^2 + \left(\frac{2}{n} \sum_{i=1}^n \cos \alpha_i \right)^2$$

V.S. Ptuskin, Adv. Space Res. 19 (1997) 697

J. Candia et al., J. Cosmol. Astropart. Phys. 5 (2003) 3

T. Antoni et al, ApJ 604 (2004) 687

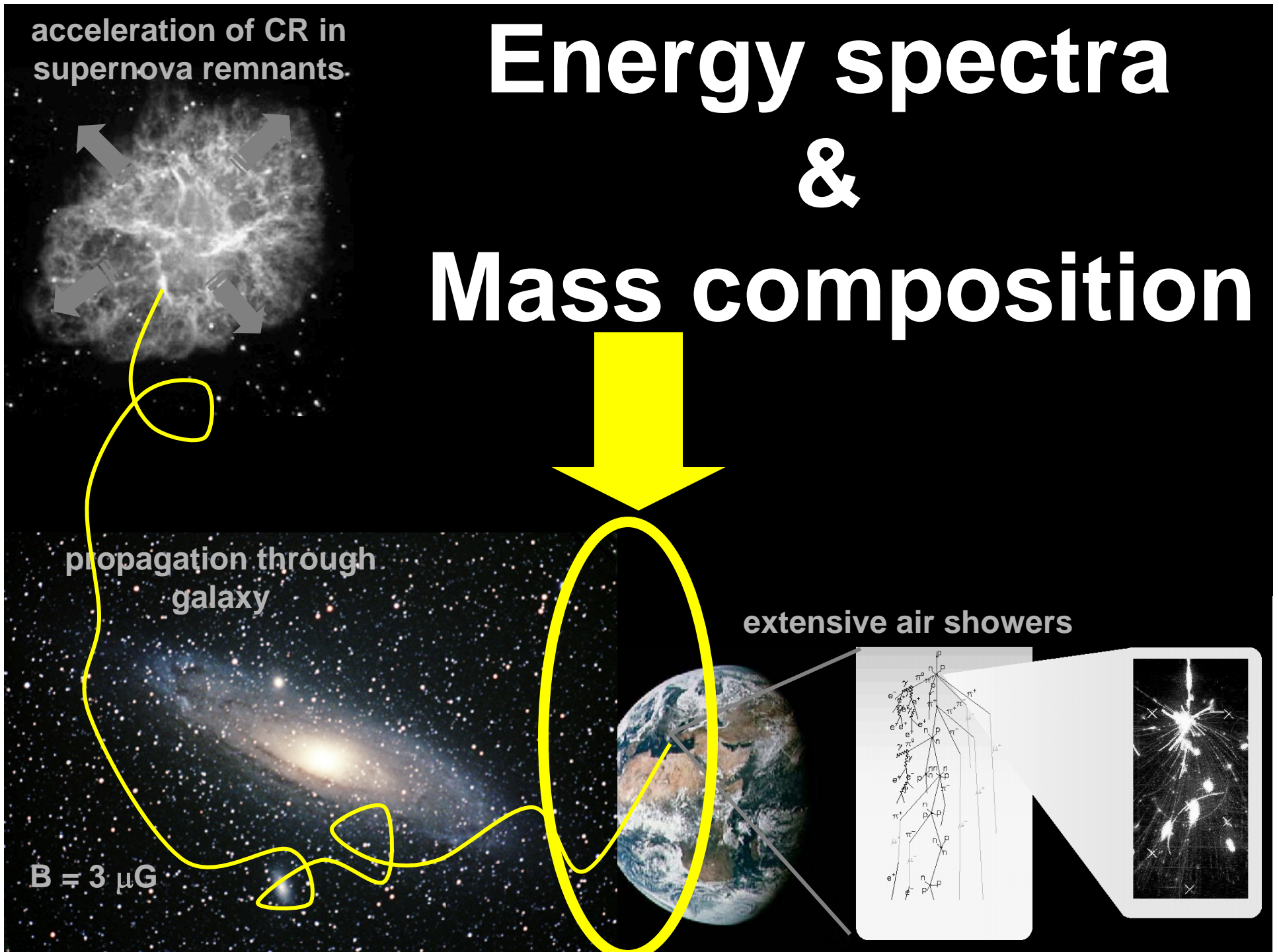
acceleration of CR in
supernova remnants

Energy spectra & Mass composition

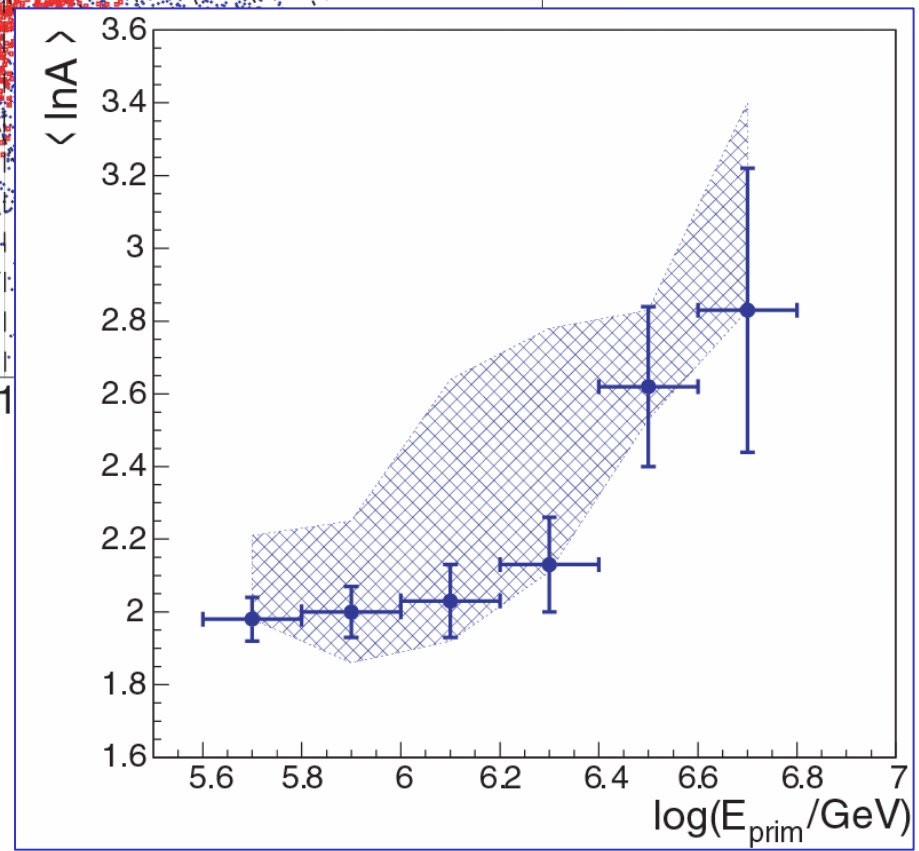
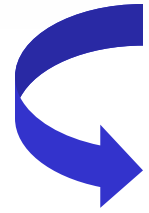
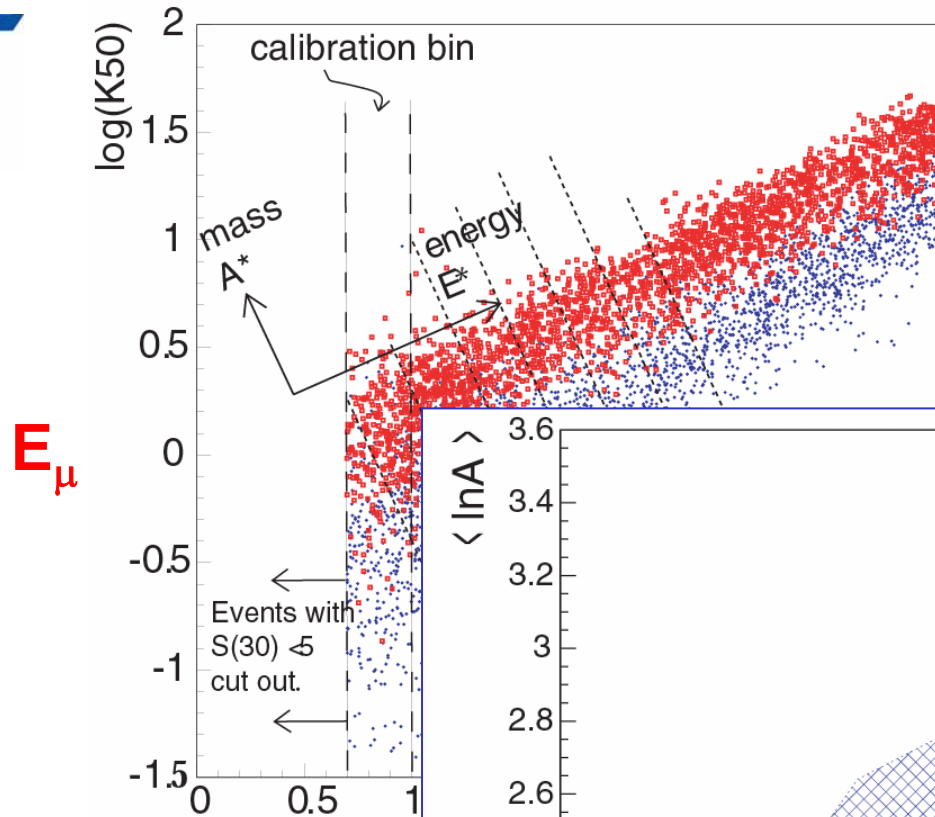
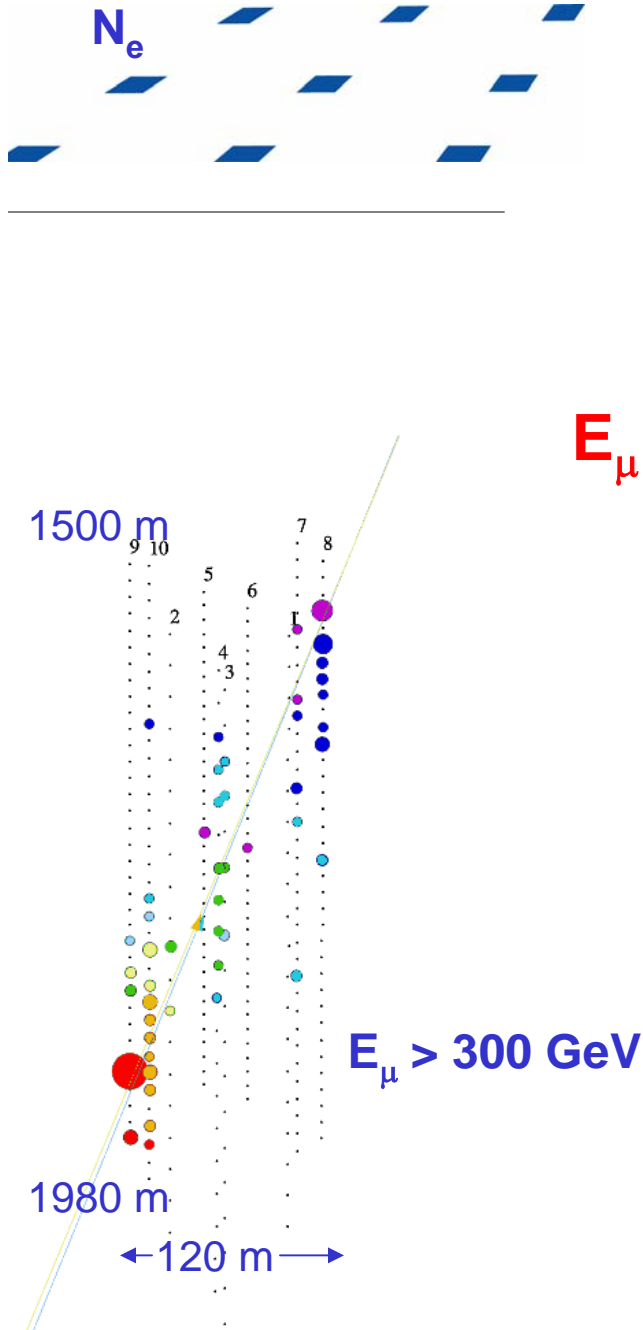
propagation through
galaxy

extensive air showers

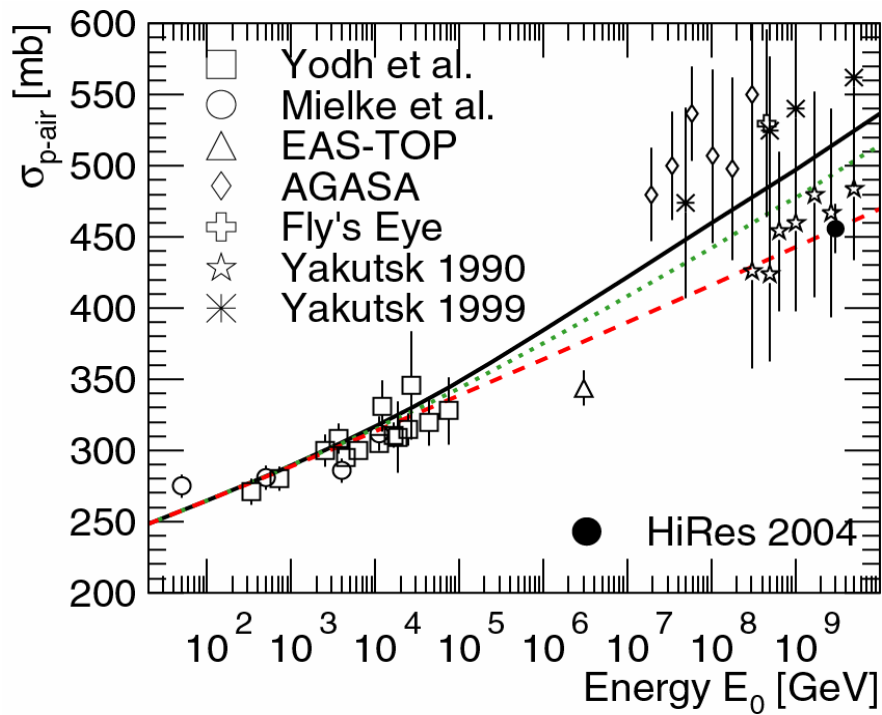
$B = 3 \mu\text{G}$



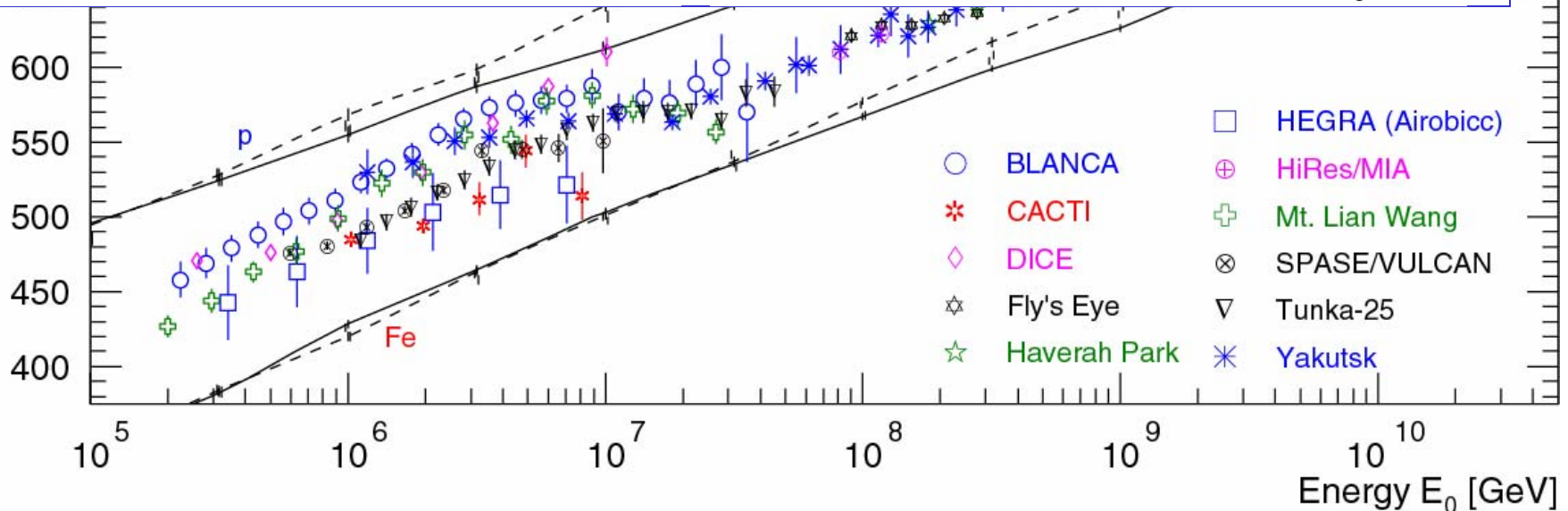
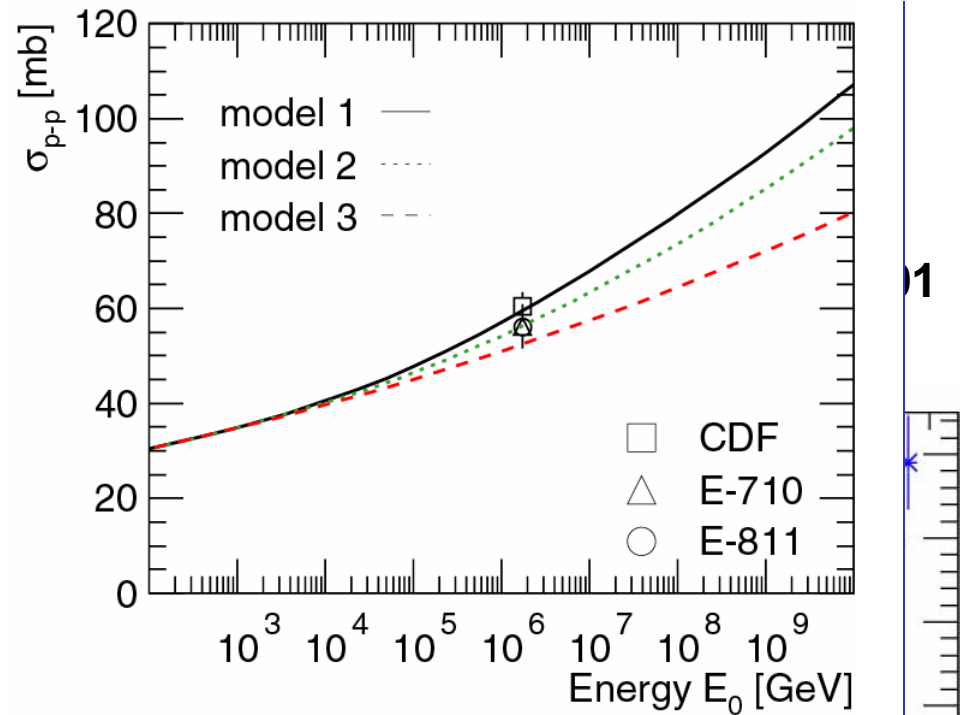
SPASE-2 / AMANDA-B10 (South Pole)



p-air cross section

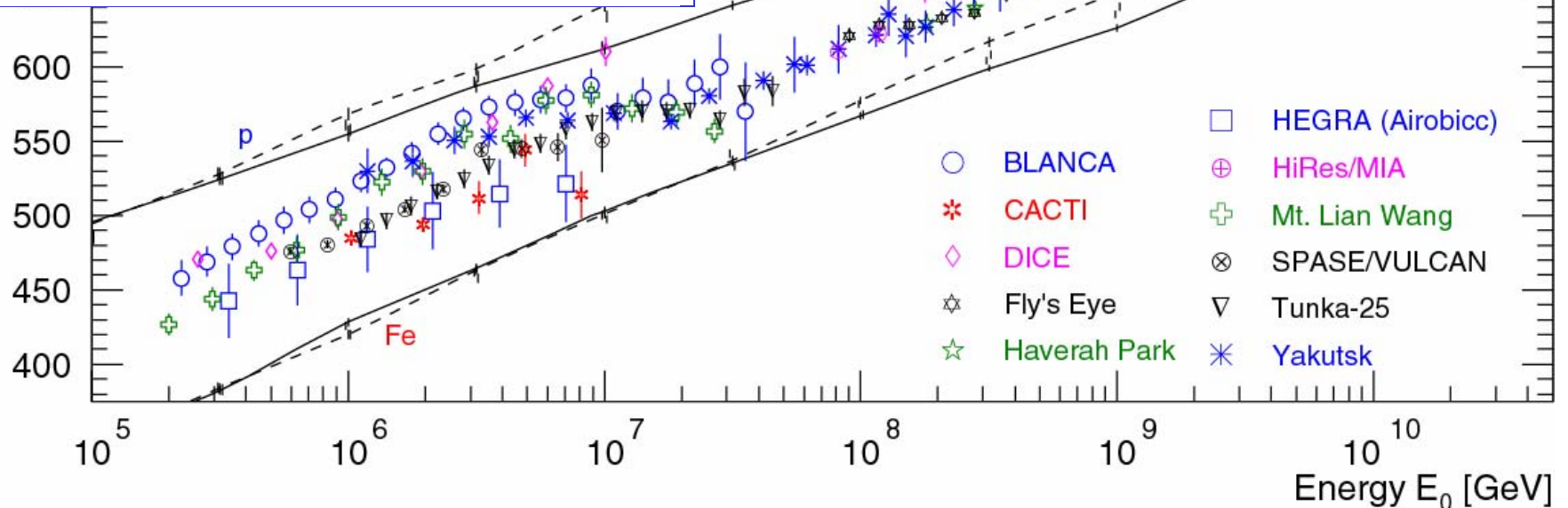
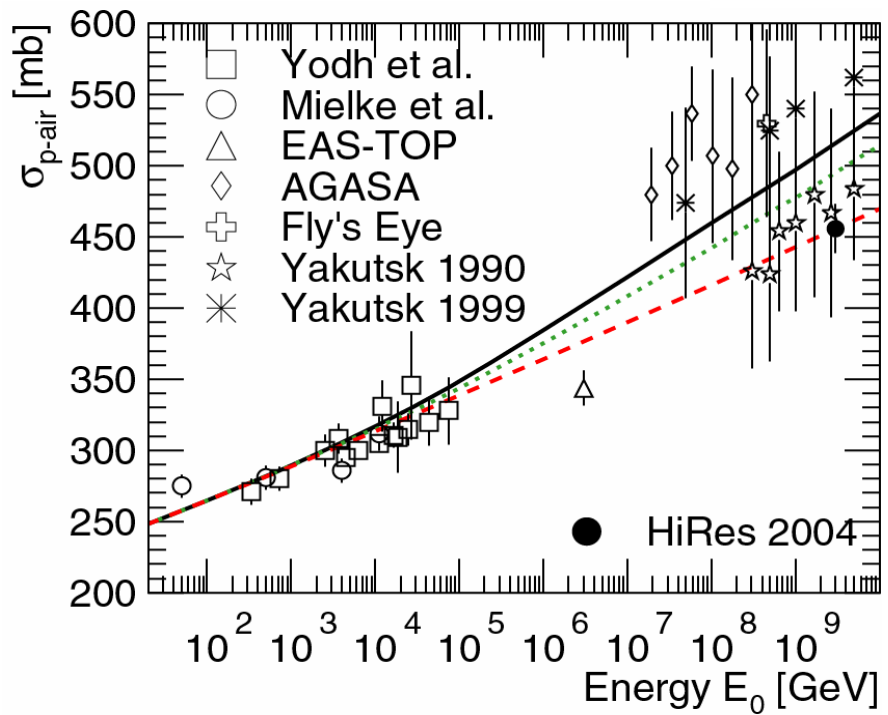


p-p cross section



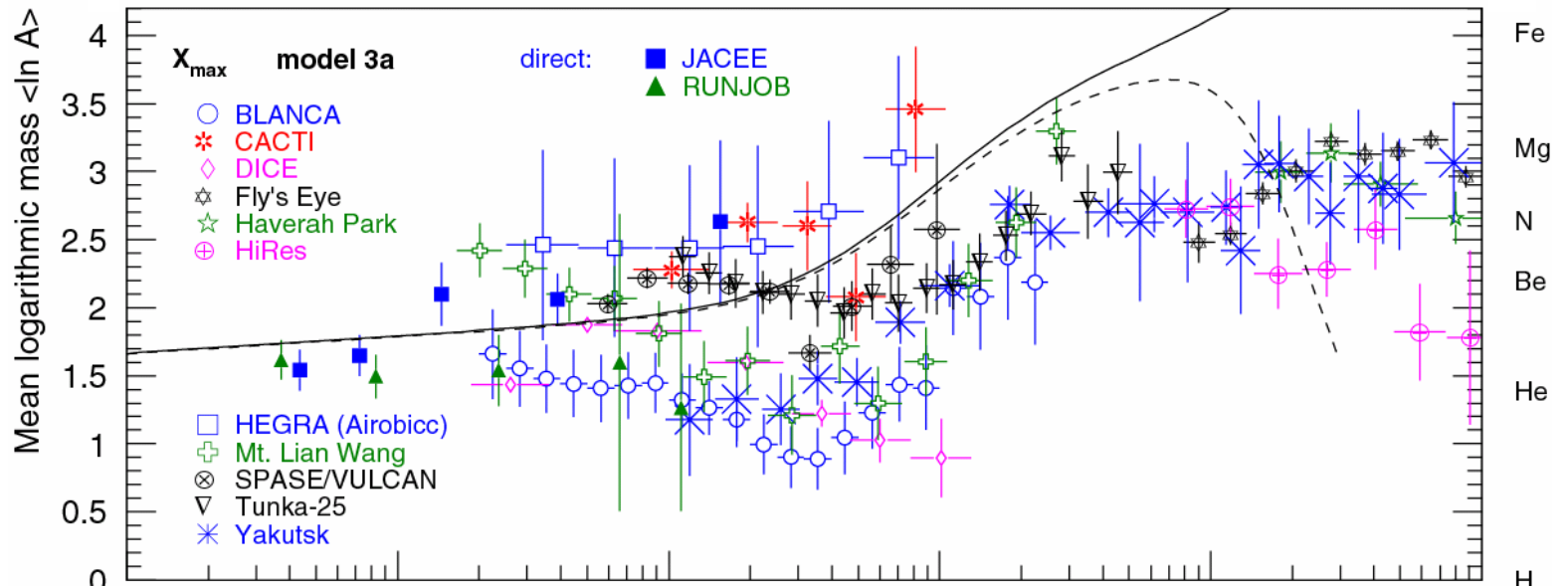
p-air cross section

Average depth of shower maximum

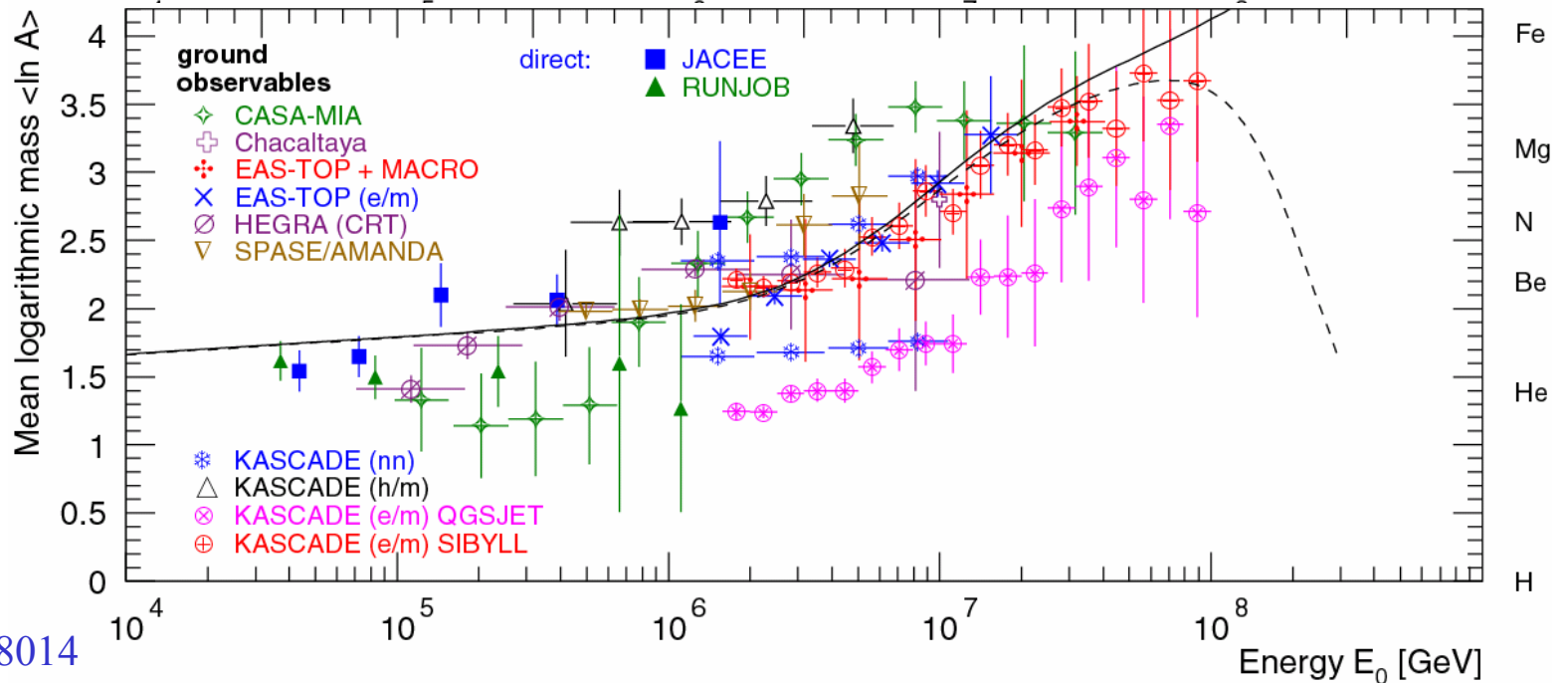


Mean logarithmic mass

X_{\max}

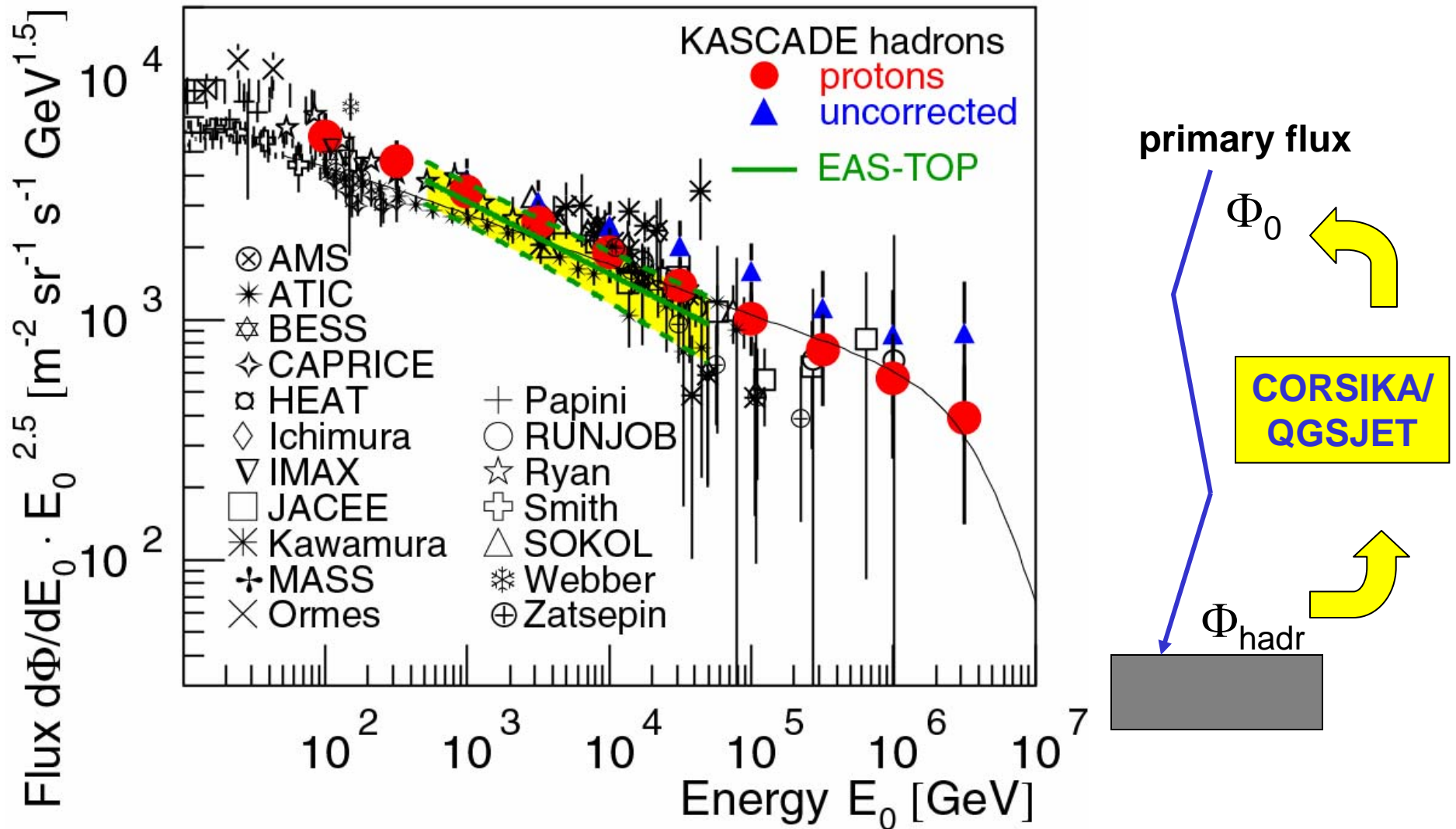


e, μ, h

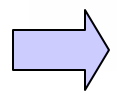
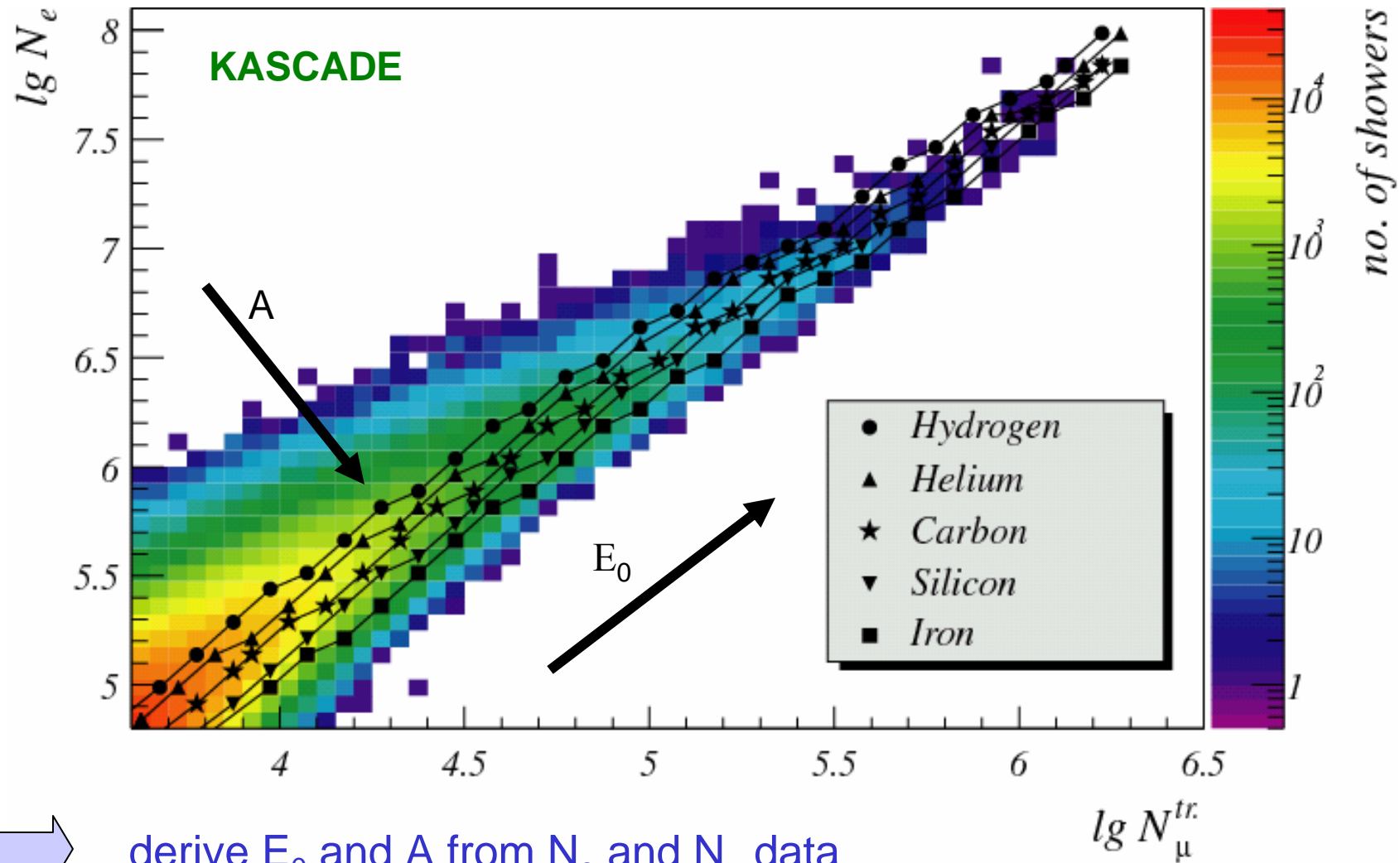


$$\langle \ln A \rangle = \sum r_i \ln A_i$$

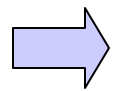
Primary proton spectrum reconstructed from unaccompanied hadrons



Two dimensional shower size spectrum $\lg N_e$ vs. $\lg N_\mu$



derive E_0 and A from N_e and N_μ data



Fredholm integral equations of 1st kind:

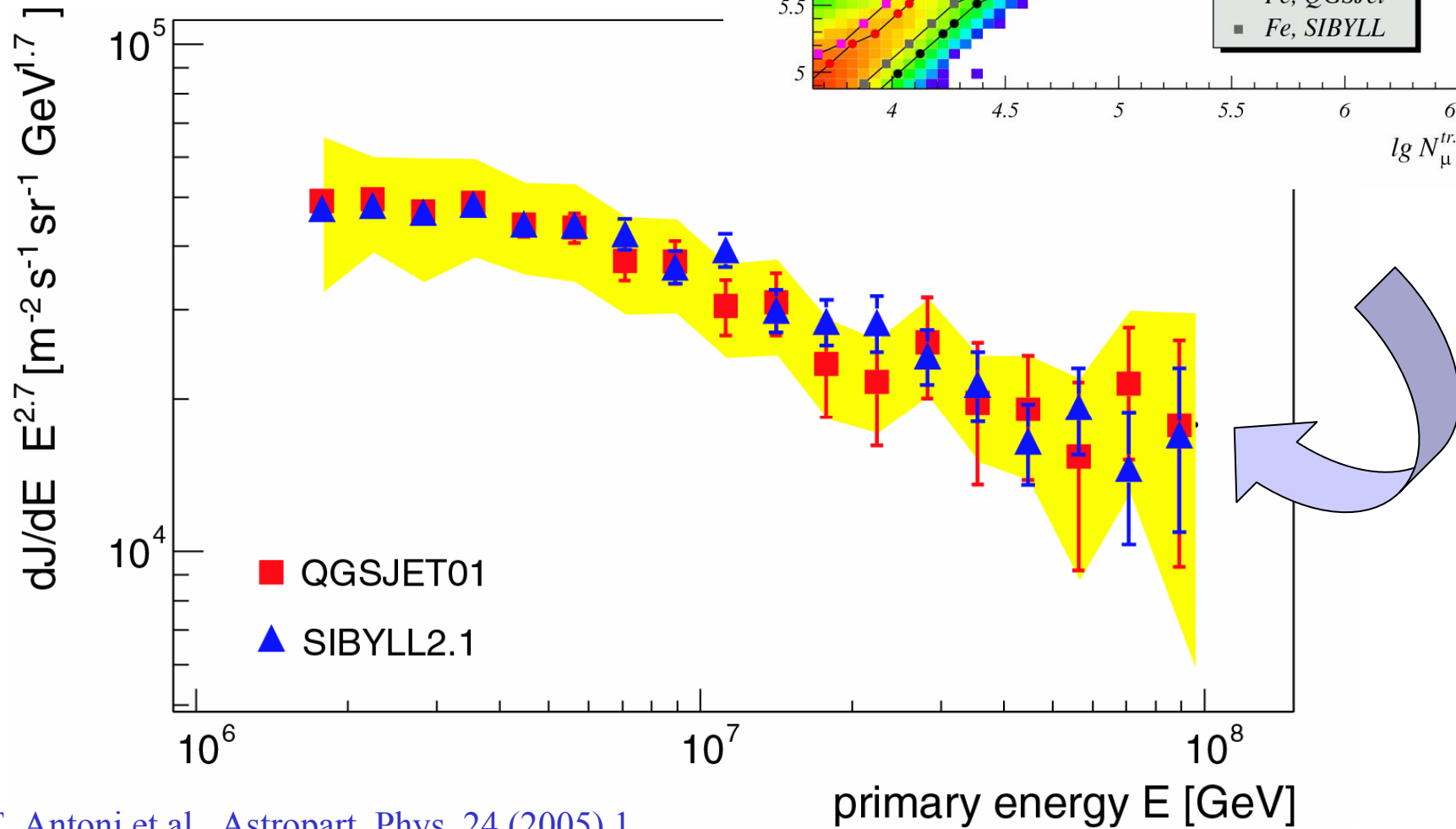
$$g_i(\lg N_e, \lg N_\mu) = \int_0^\infty t_i(\lg N_e, \lg N_\mu | E) p_i(E) dE$$

All-particle energy spectrum

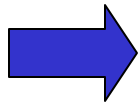
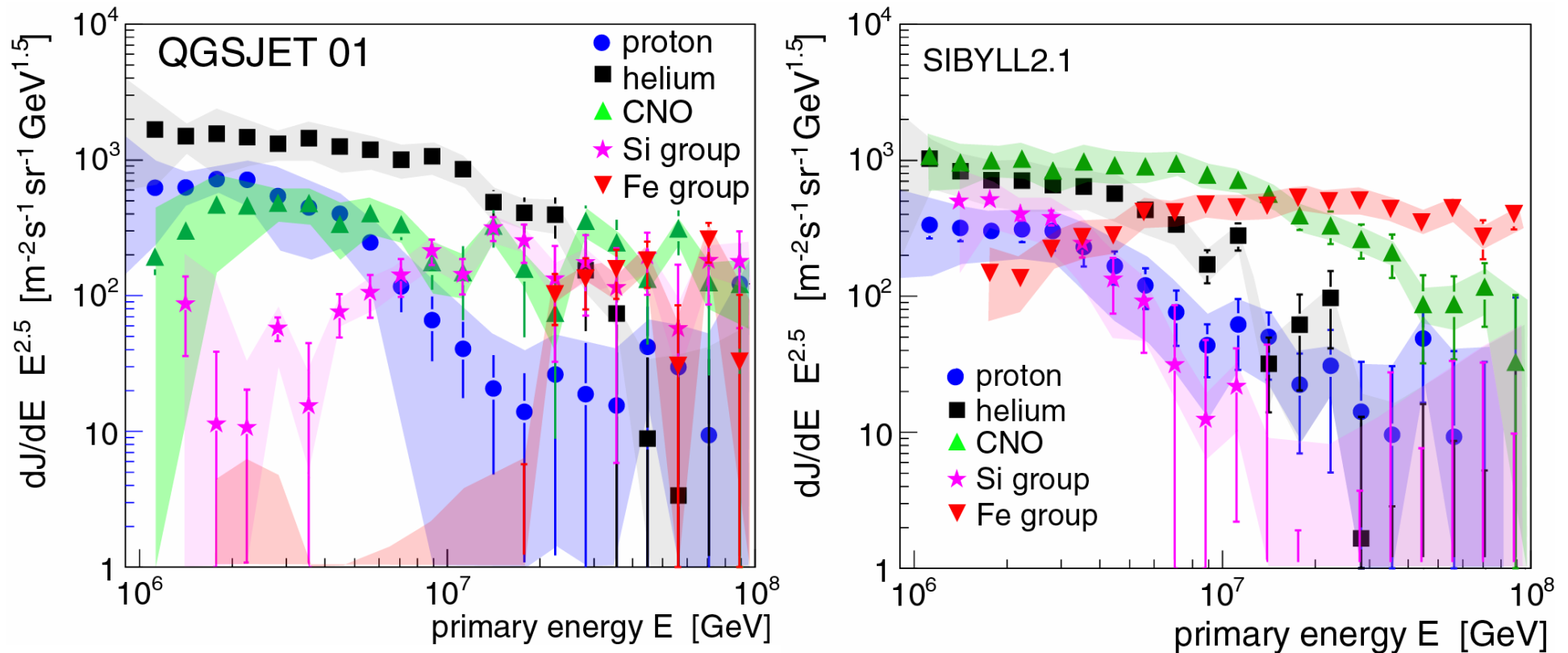
two hadronic interaction models:
CORSIKA 6.018/GHEISHA 2002

- QGSJET 01

- SIBYLL 2.1

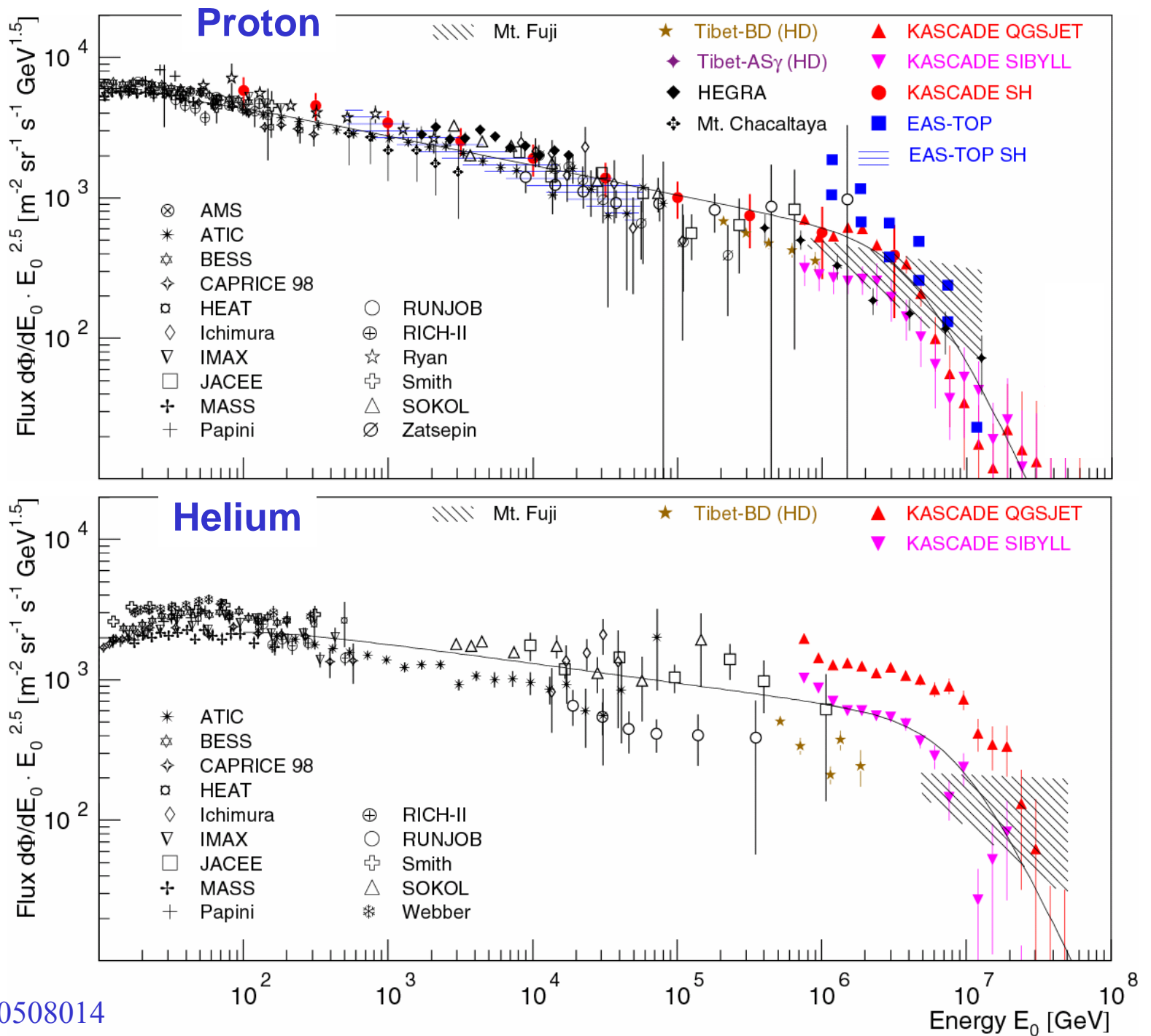


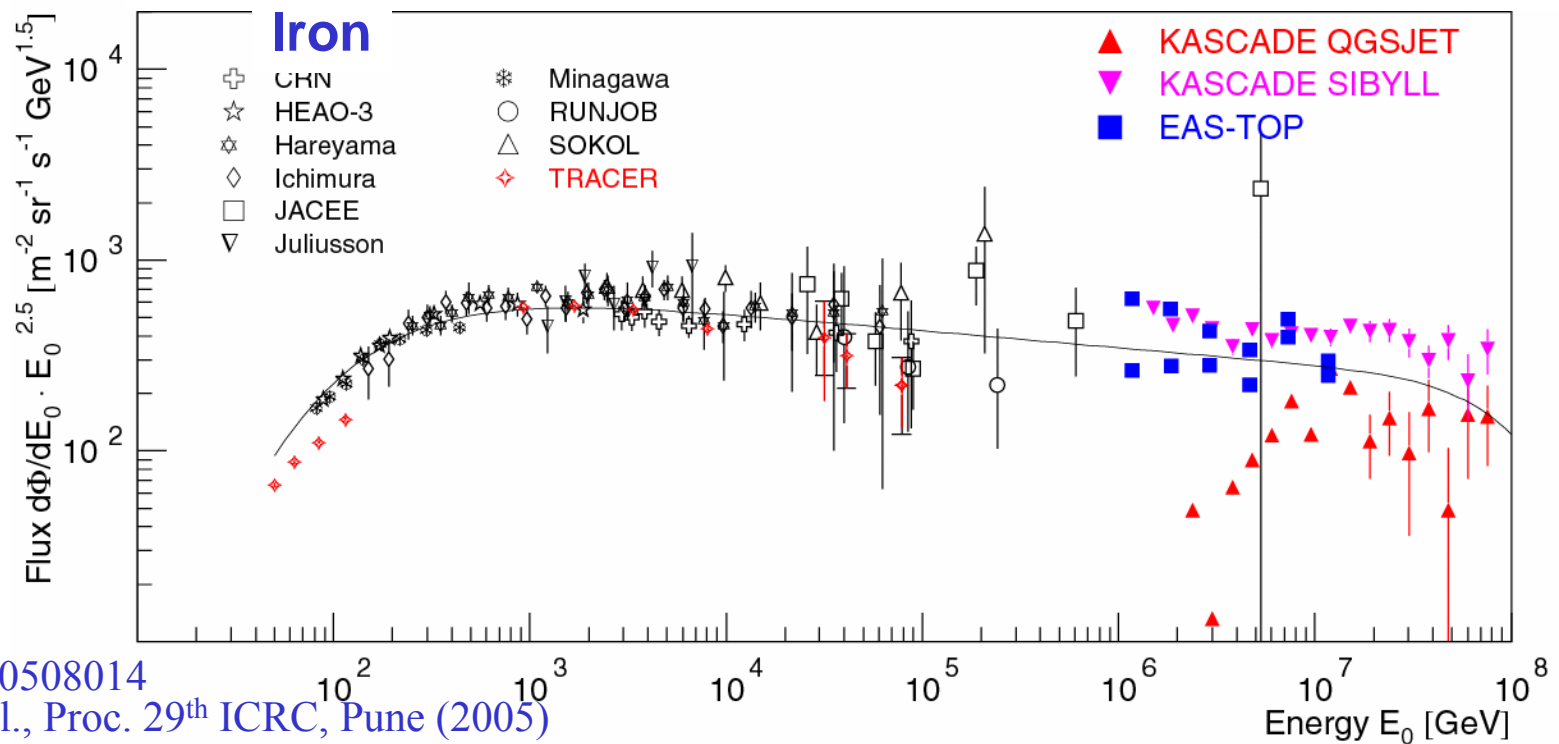
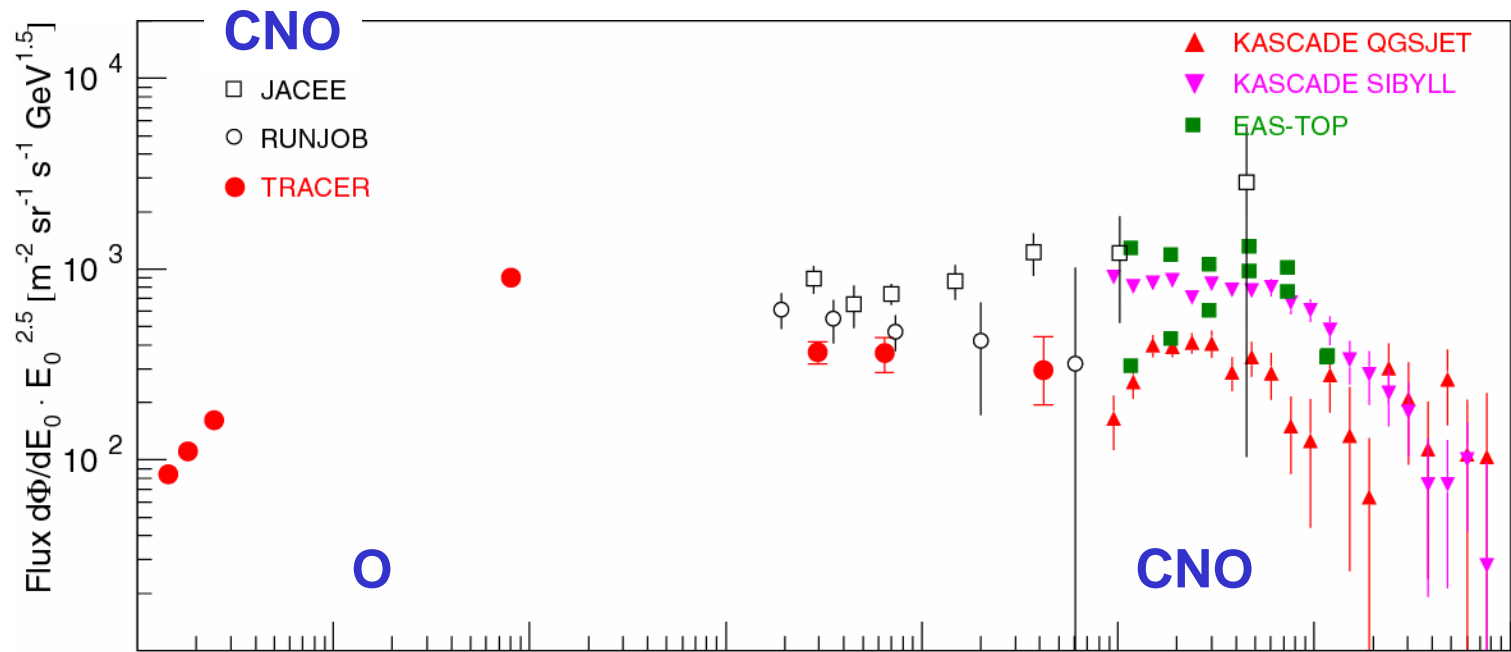
KASCADE: Energy spectra for elemental groups



Knee caused by cut-off for light elements

Astrophysical interpretation limited by description of interactions in the atmosphere





KASCADE GRANDE Array

37 detector stations

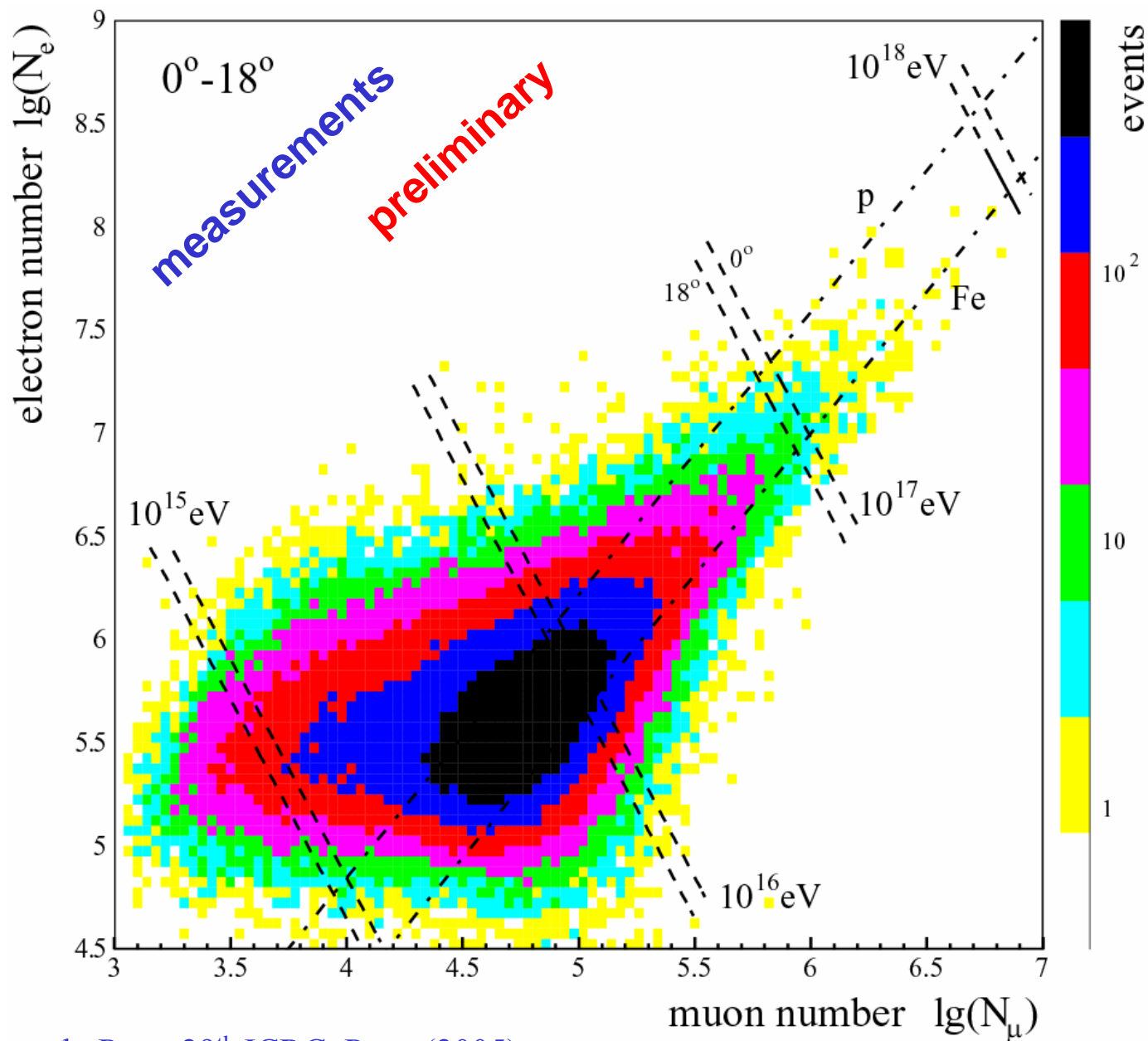
370 m² e/γ:
Scintillation counter



KASCADE
100 m x 200 m

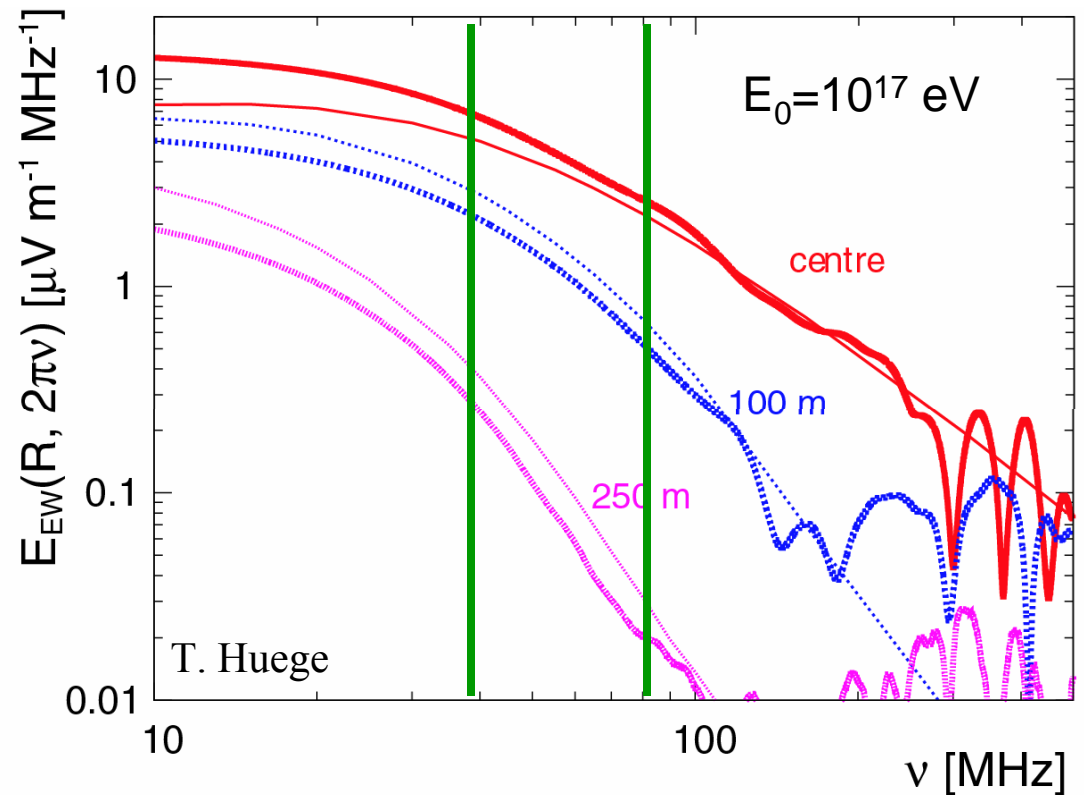
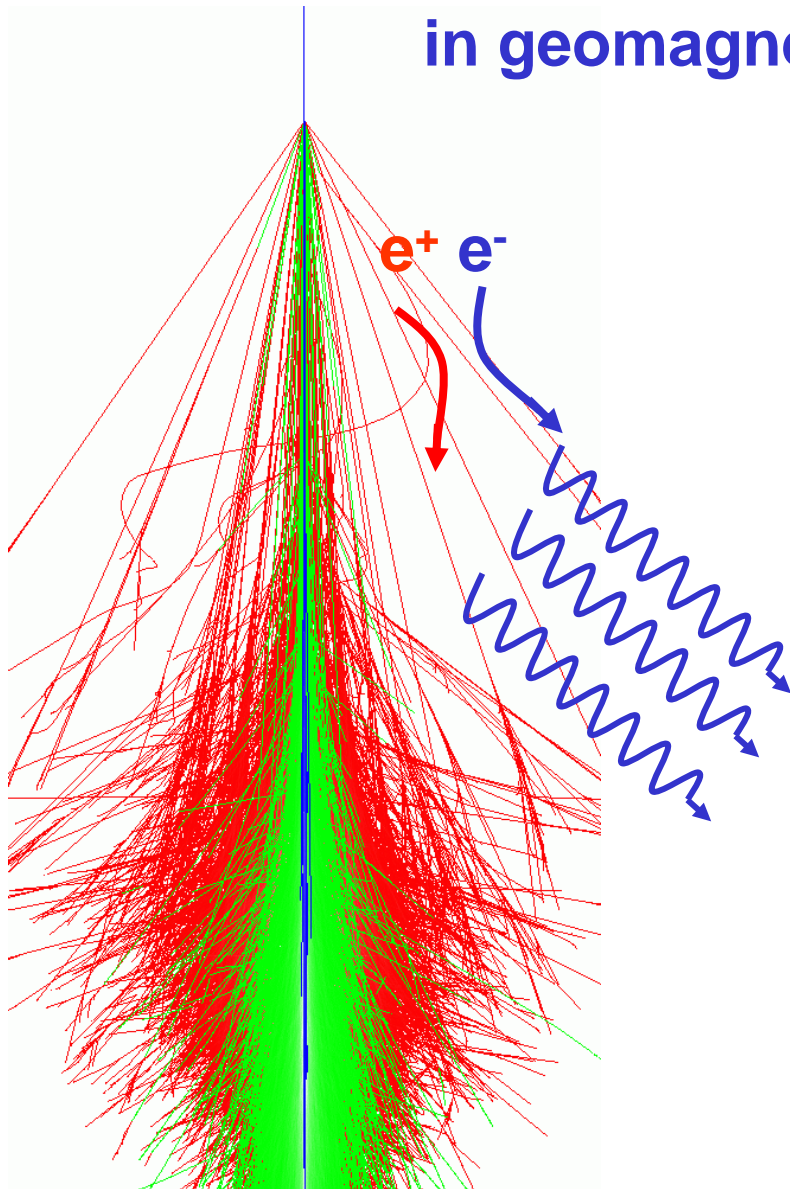


KASCADE-Grande – N_e - N_μ correlation



Radio emission from air showers - LOPES

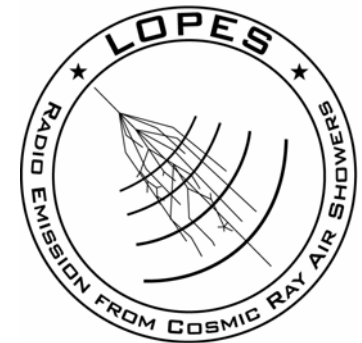
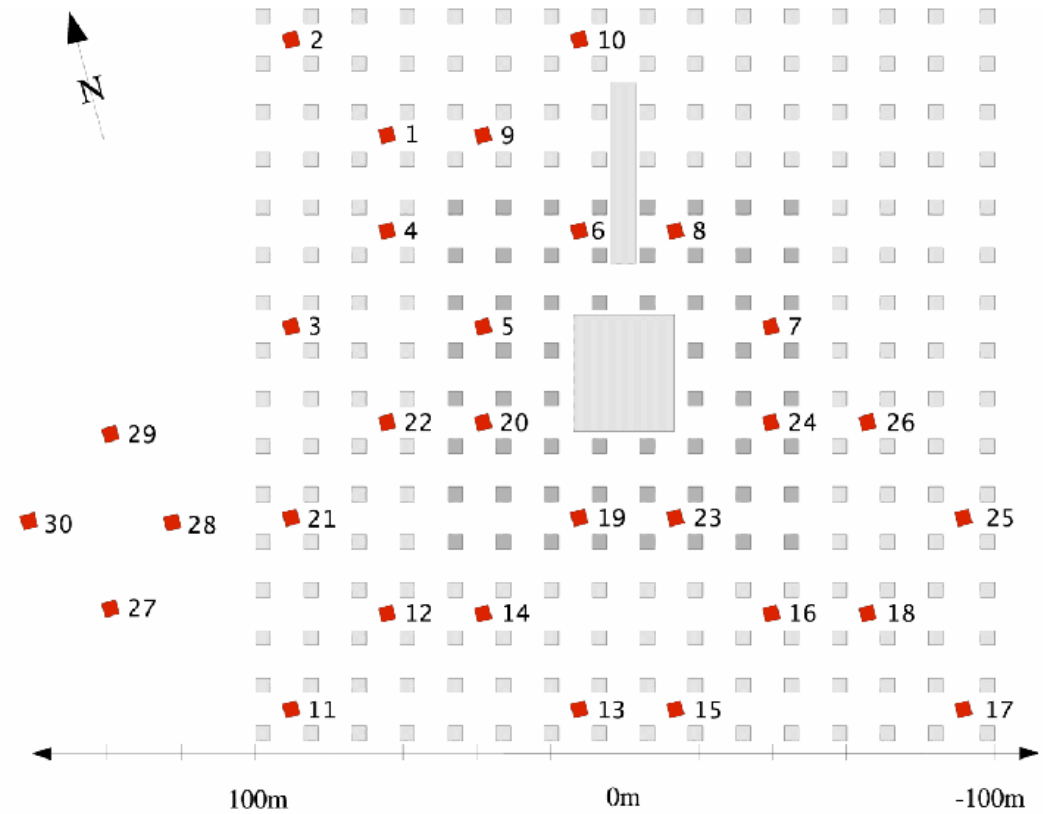
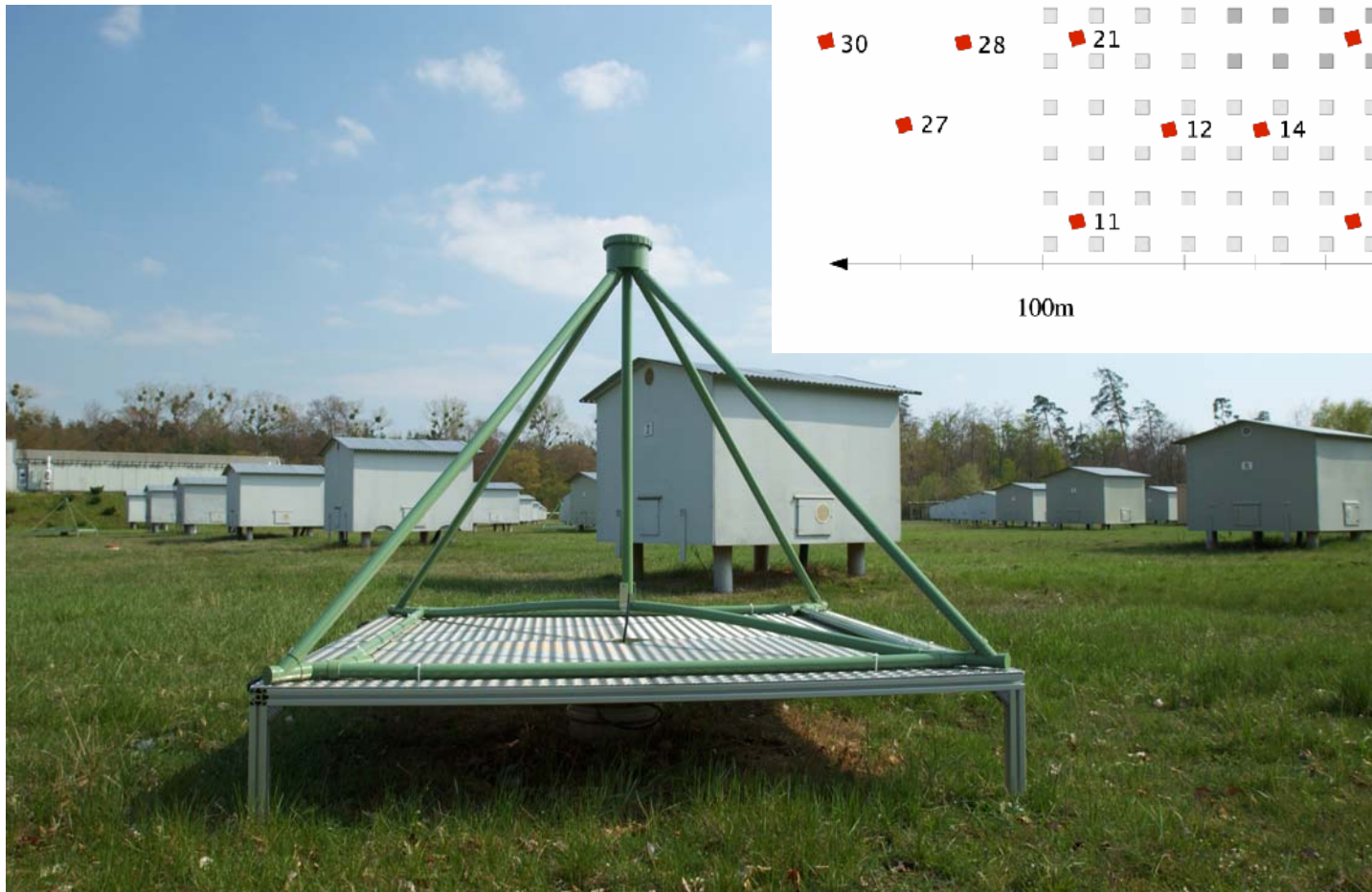
Coherent emission of synchrotron radiation
in geomagnetic field



expect signals $\sim \mu\text{V}/(\text{m MHz})$

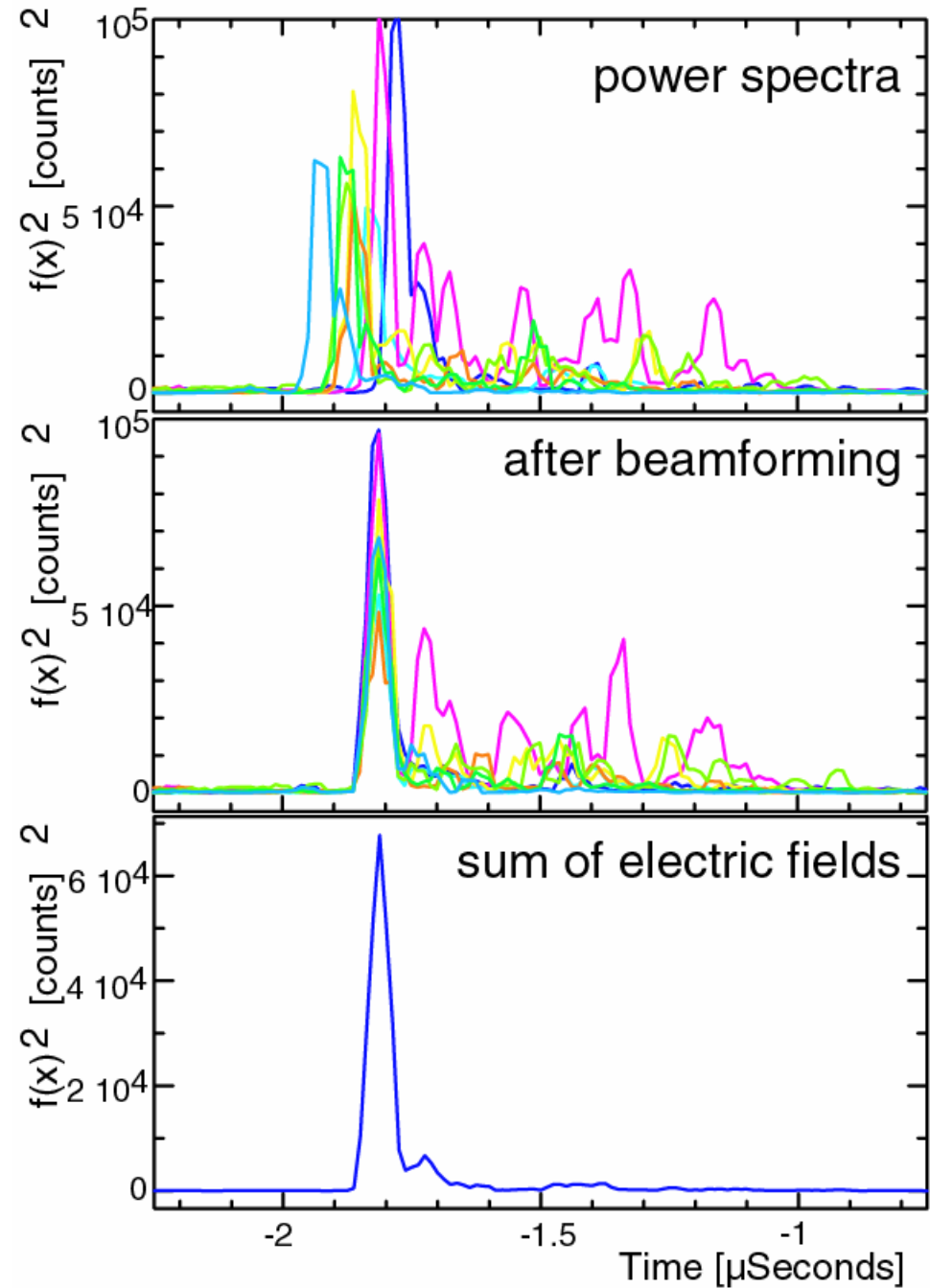
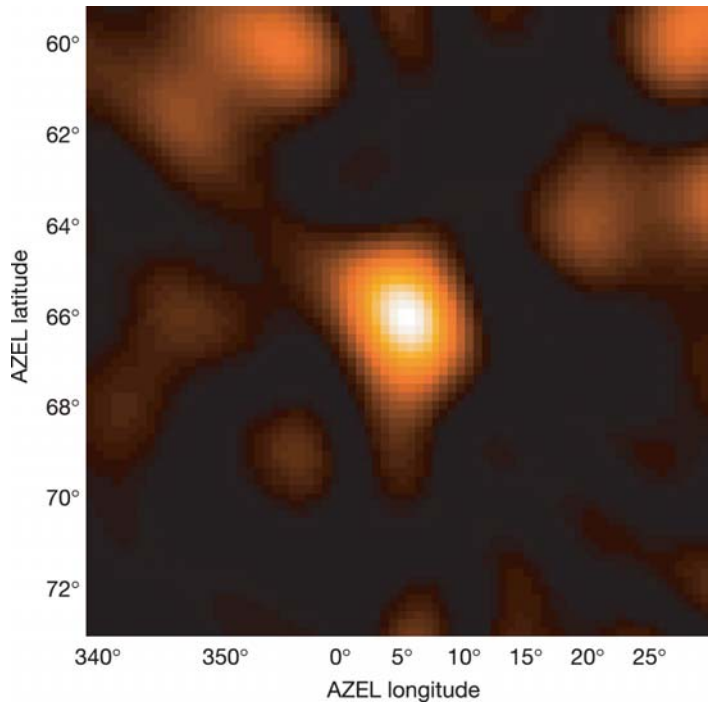
LOPES

30 antennas operating at
KASCADE-Grande



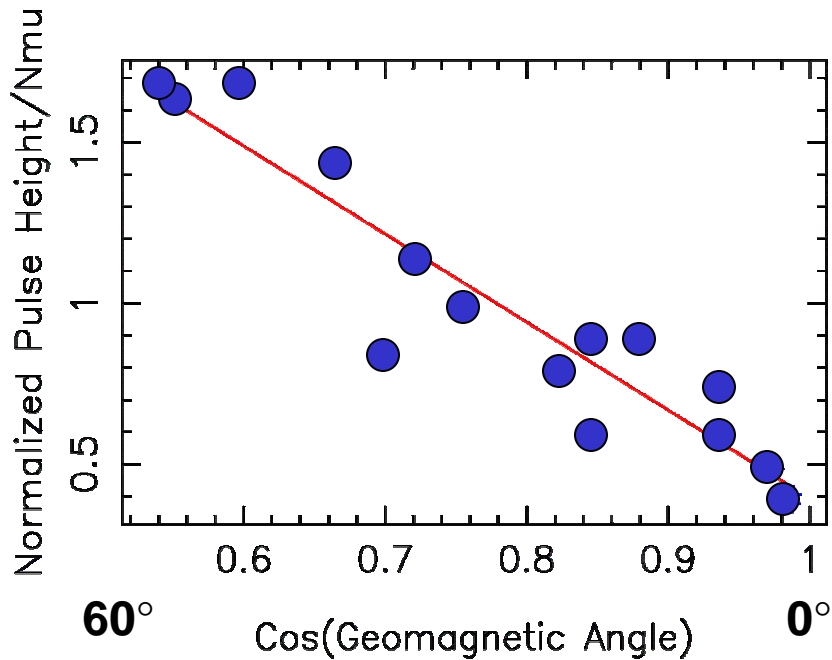
LOPES first signals

Position of shower in sky

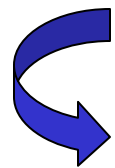
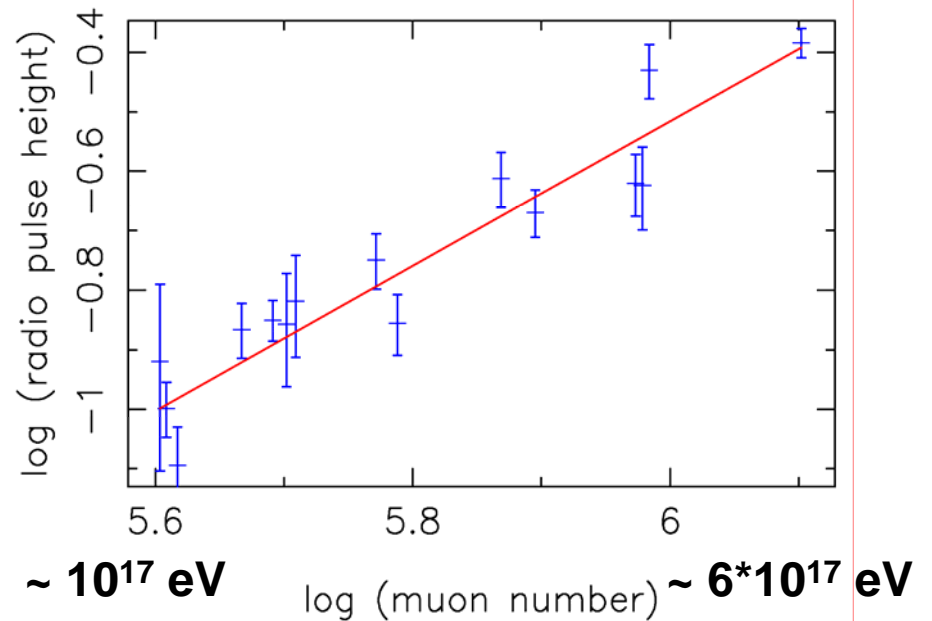


Radio signal – dependence on

angle with respect to
geomagnetic field



number of muons
(i.e. primary energy)



Geosynchrotron emission



**Radio signal
increases with energy**

Hadrons at high altitude → surviving protons

calorimeter @ 500 g/cm²

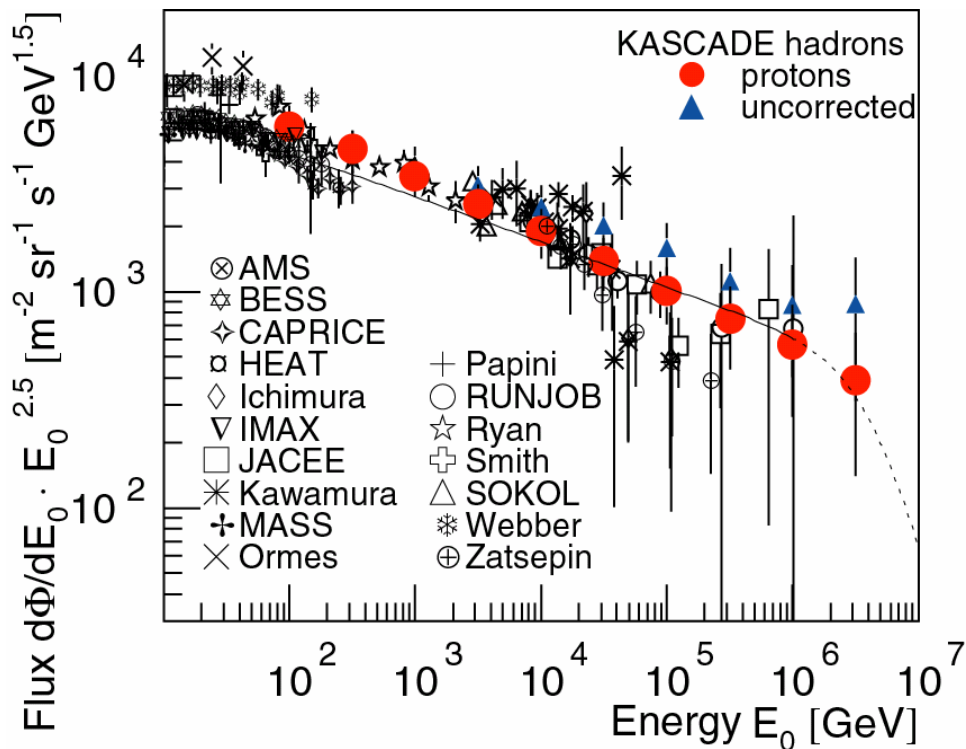
1 PeV: ~ 6.5 λ_i

320 m² sr



0.5 m² sr effective

ideal: combination with air Cerenkov detector for calibration



From the Knee to the toes: The challenge of cosmic-ray composition

Status:

- Description of interactions in atmosphere improved
- Mean mass increases as function of energy (knee region)
- Knee is caused by subsequential cut-offs for individual elements
- Astrophysical interpretation of EAS measurements limited by understanding of interactions in the atmosphere

Perspectives:

- KASCADE-Grande (2nd knee)
- LOPES (radio signals)

} galactic → extragalactic
10¹⁸ eV

- Radio in AUGER

>10²⁰ eV

- Hadrons at large altitude → Proton spectrum

} direct ↔ indirect
10¹⁵ eV

Thanks to my
colleagues ...

**KASCADE-Grande
Collaboration**

Experimentelle Teilchenphysik,
University of Siegen

M. Brüggemann, P. Buchholz,
C. Grupen, Y. Kolotaev,
W. Walkowiak, D. Zimmermann

Institut für Kernphysik
Forschungszentrum and University of Karlsruhe

T. Antoni, W.D. Apel, F. Badea, K. Bekk, A. Bercuci, H. Blümer,
H. Bozdog, C. Büttner, K. Daumiller, P. Doll, R. Engel, J. Engler,
F. Feßler, H.J. Gils, A. Haungs, D. Heck, J.R. Hörandel,
H.O. Klages, G. Maier, H.-J. Mathes, H.J. Mayer, J. Milke,
M. Müller, R. Obenland, J. Oehlschläger, S. Ostapchenko,
S. Plewnia, H. Rebel, M. Roth, G. Schatz, H. Schieler, J. Scholz,
M. Stümpert, T. Thouw, H. Ulrich, J. van Buren, A. Weindel,
J. Wochele, S. Zagrömski

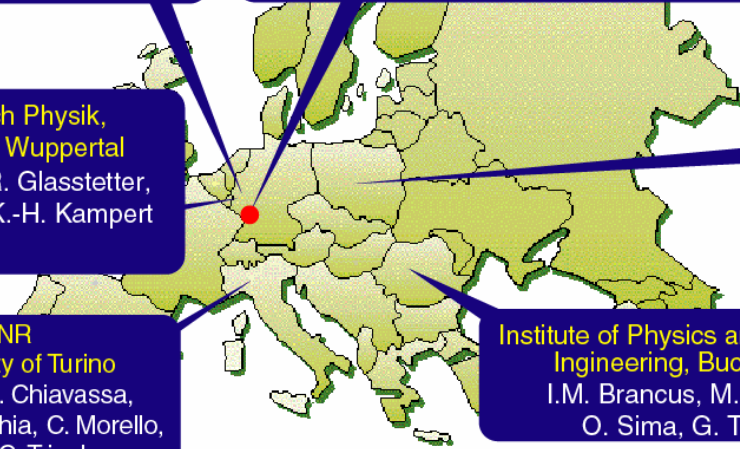
Fachbereich Physik,
University of Wuppertal

K.-H. Becker, R. Glasstetter,
S. Hartmann, K.-H. Kampert

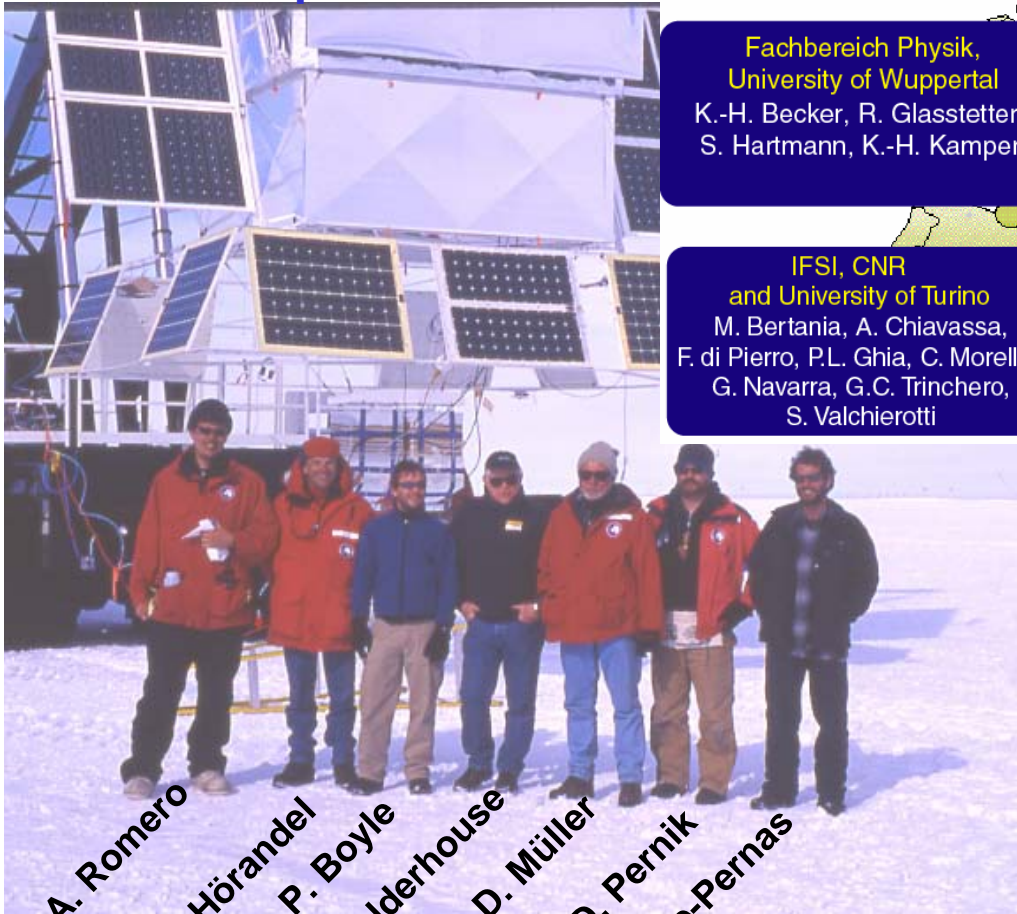
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